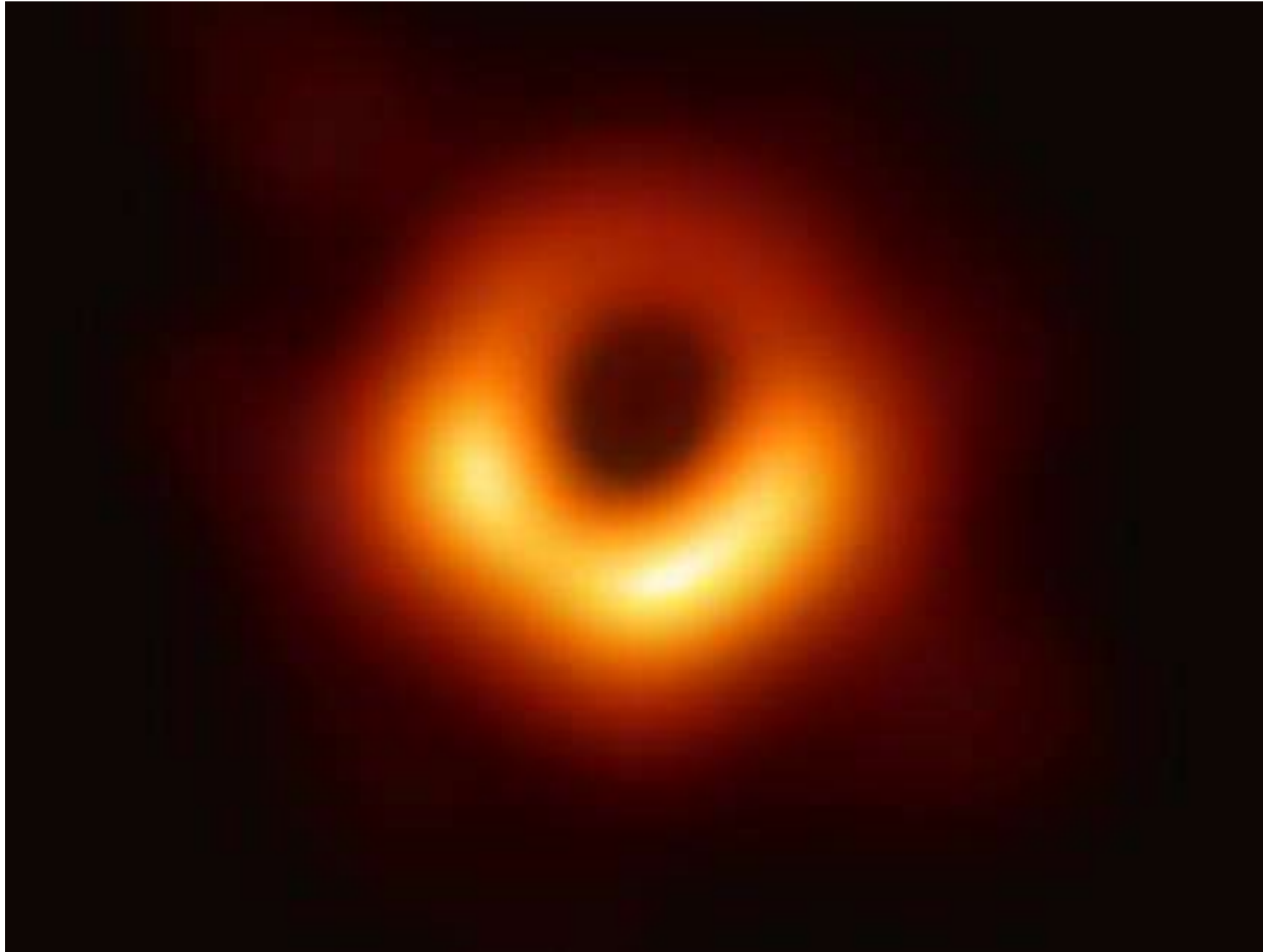


Hogyan készült kép a fekete lyuk árnyékáról?

Gabányi Krisztina (ELTE TTK Csillagászati Tanszék, MTA-ELTE
Extragalaktikus Asztrofizika Kutatócsoport, Csillagászati és
Földtudományi Központ Konkoly-Thege Miklós Csillagászati Intézet)

Event Horizon Telescope

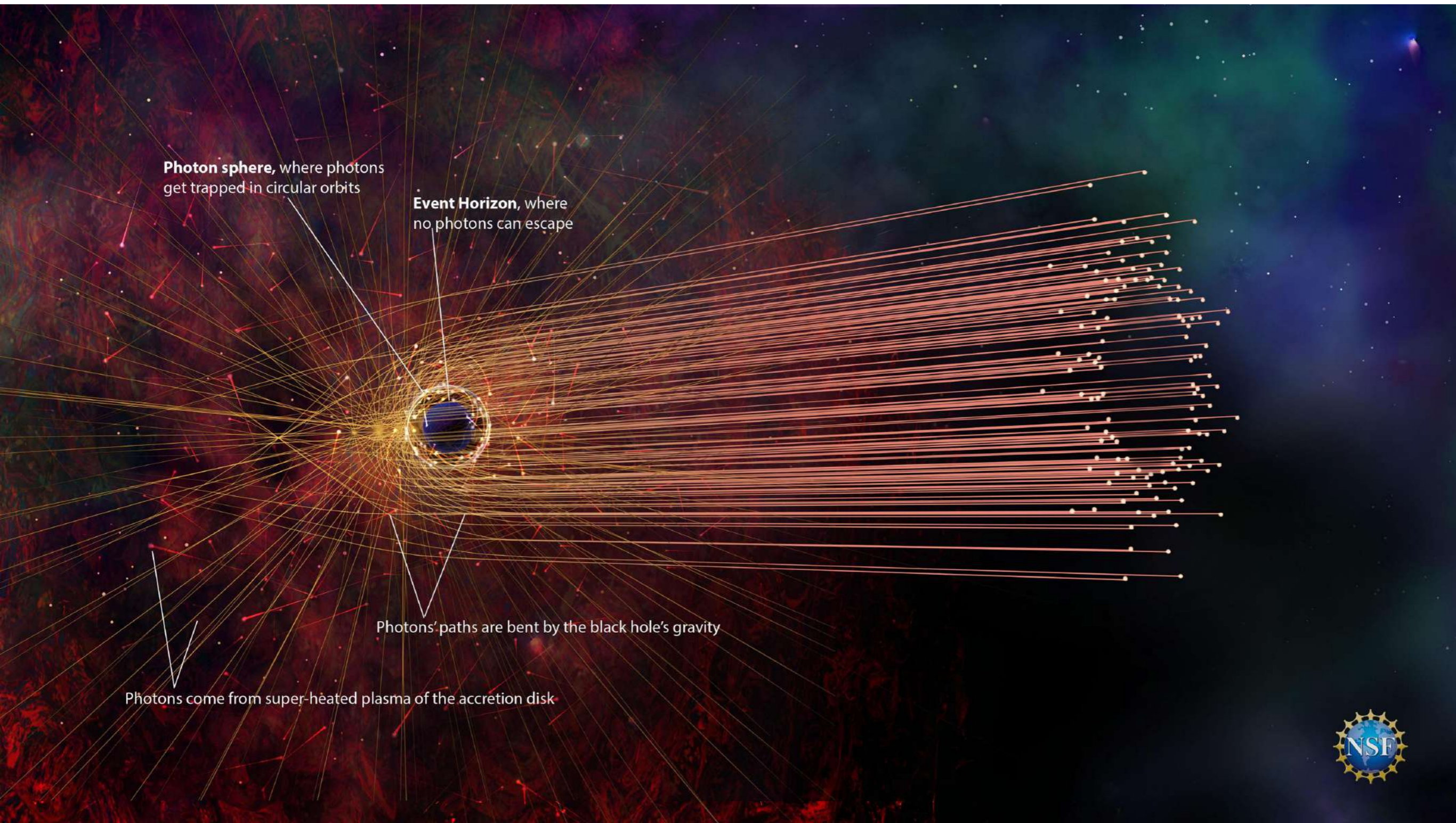
6,5 milliárd naptömegű
fekete lyuk

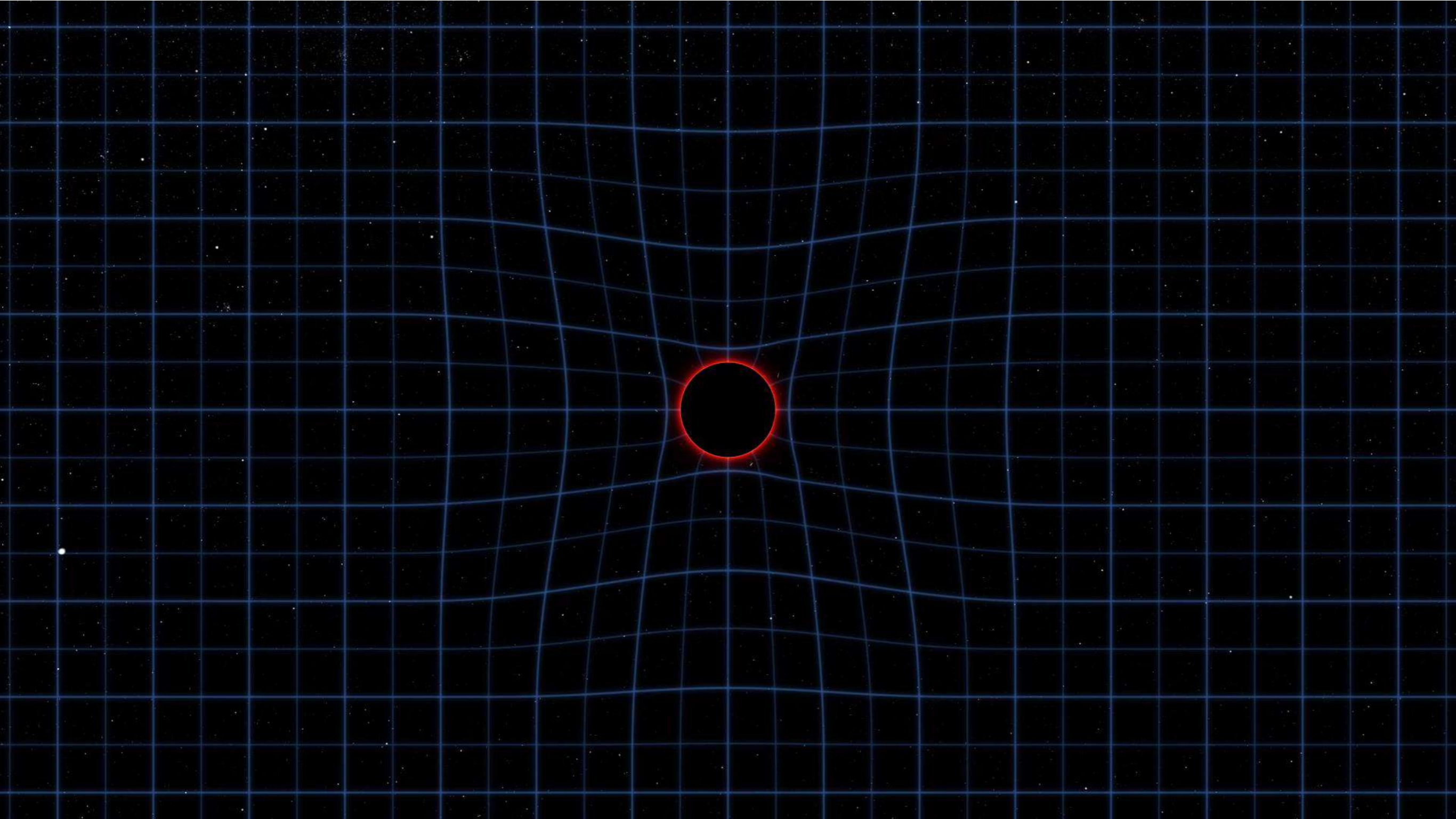


Kép egy fekete lyukról?

- Miről tudunk képet alkotni, ha ez egy fekete lyuk?
- Hogyan tudunk képet alkotni valamiről, ami ennyire kicsi?

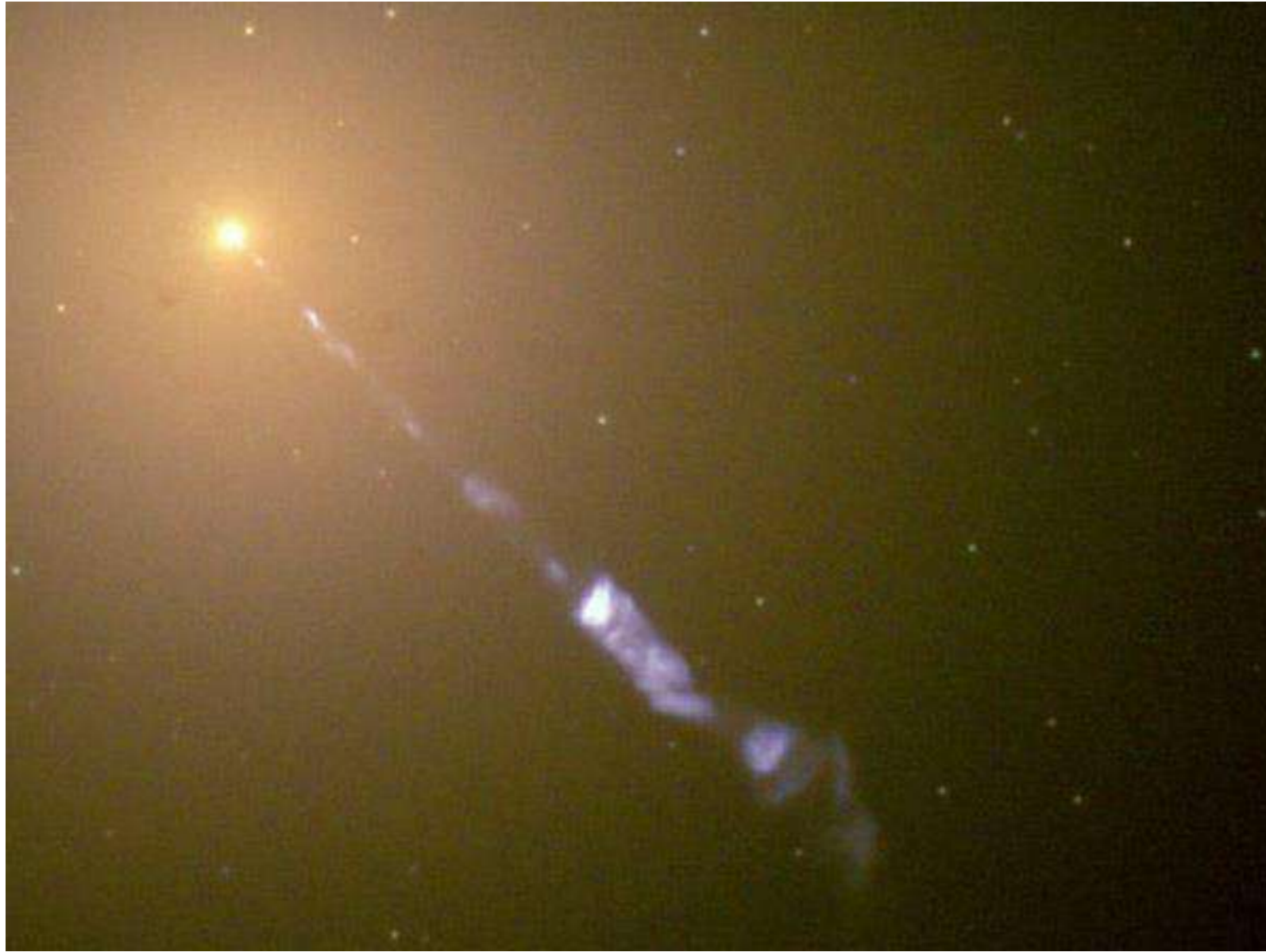
Miről alkotunk képet?





Hubble-űrtávcső felvétele

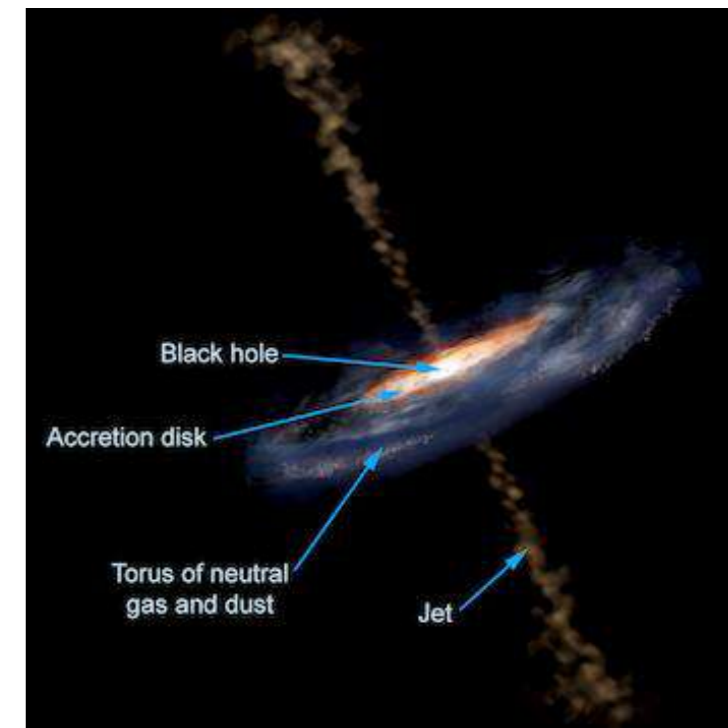
M87



55 millió fényévre
Óriás elliptikus galaxis a Virgo
halmazban
1918 - H.D. Curtis: „furcsa
egyenes nyaláb”
1950: Virgo A az egyik
legfényesebb rádióforrás az
égen
6,5 milliárd naptömegű fekete
lyuk

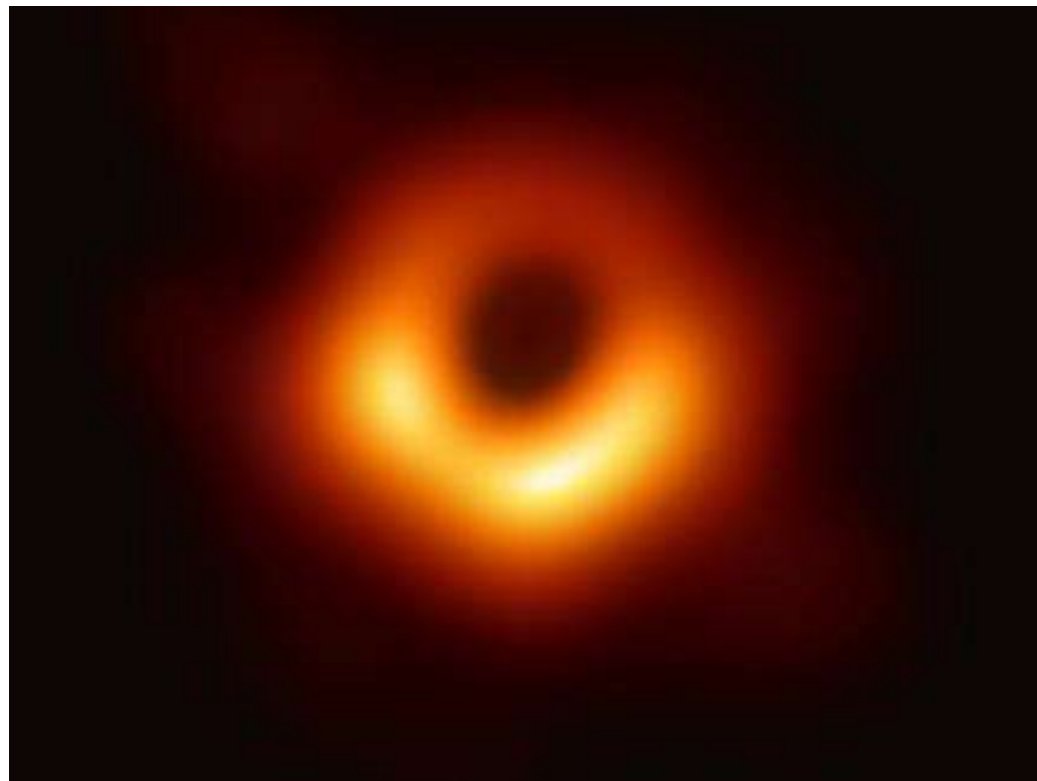
31 ívmásodperc \approx 8000 fényév

Az eseményhorizont sugara \approx 127 cs.e.





Összehasonlításképp: A Voyager-1 űrszonda 150,5 cs.e. távolságban van a Földtől jelenleg



Átmérő: $\sim 5 \frac{2GM}{c^2}$

Schwarzschild-sugár

*55 millió fényév távolságban,
40 milliomed ívmásodperc
szögméretnek felel meg*

*Ennél nagyobb szögméretű ismert fekete lyuk árnyék csak a
Tejútrendszer közepén található fekete lyukhoz tartozik:
Sagittarius A* , ≈ 4 millió naptömeg, ≈ 26700 fényévre a Földtől*

*A Hold szögmérete 0,5 fok
(45 milliószor nagyobb)*



Felbontóképesség

$$\theta \sim \frac{\lambda}{D}$$

hullámhossz

átmérő



(a)



(b)



(c)



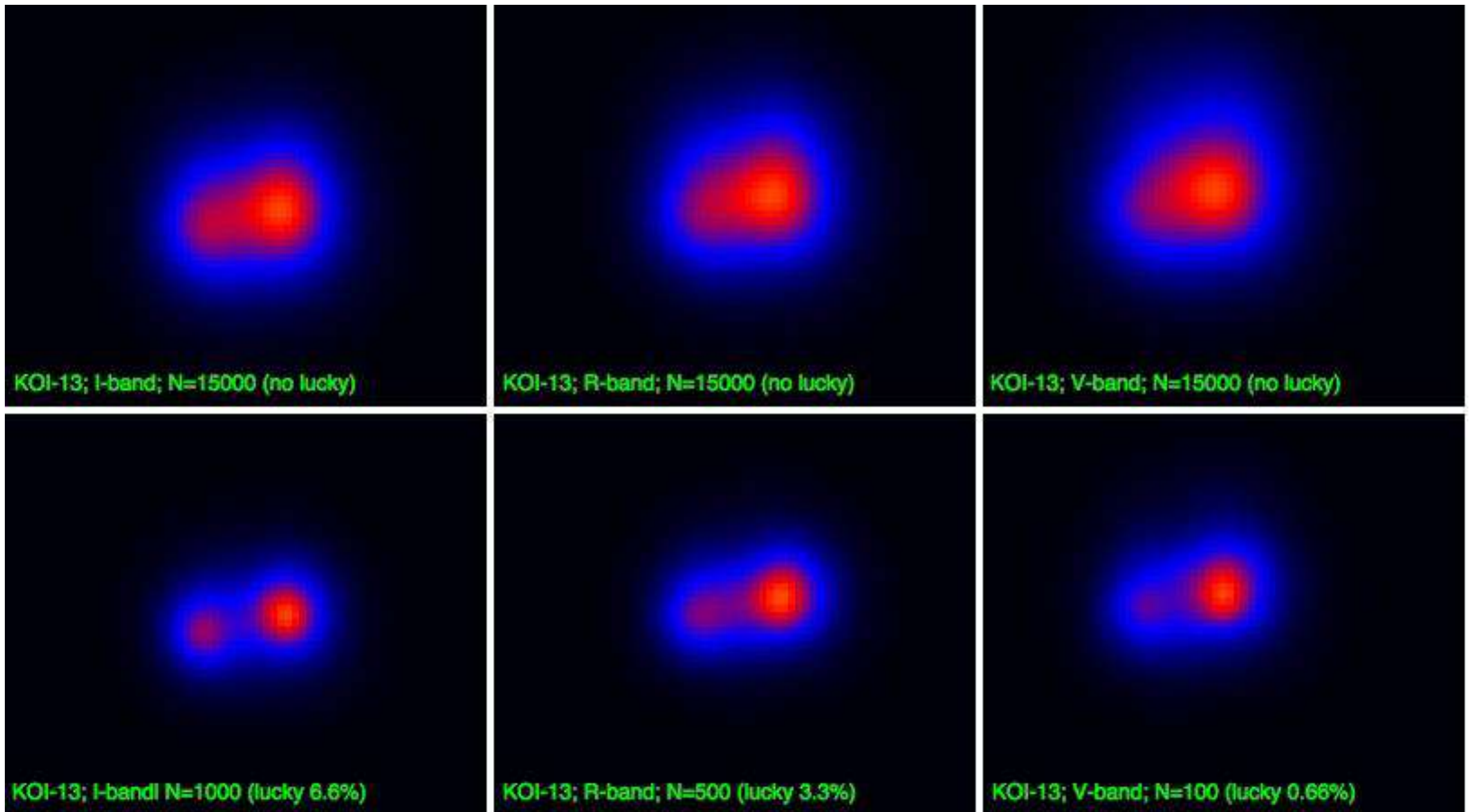
(d)

Felbontóképeség

$$\theta \sim \frac{\lambda}{D}$$

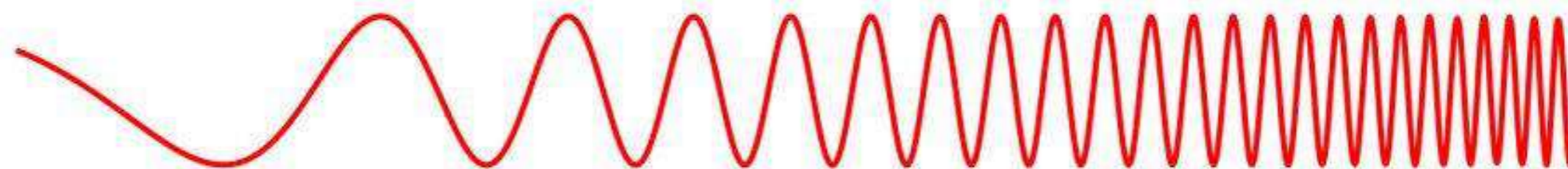
hullámhossz

átmérő

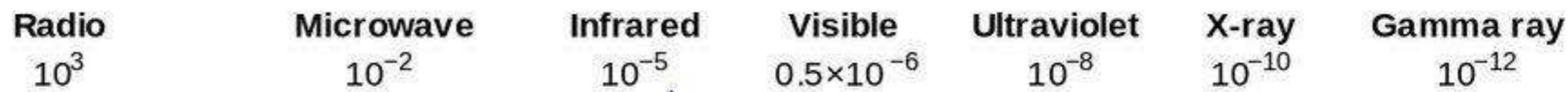


Hullámhossz

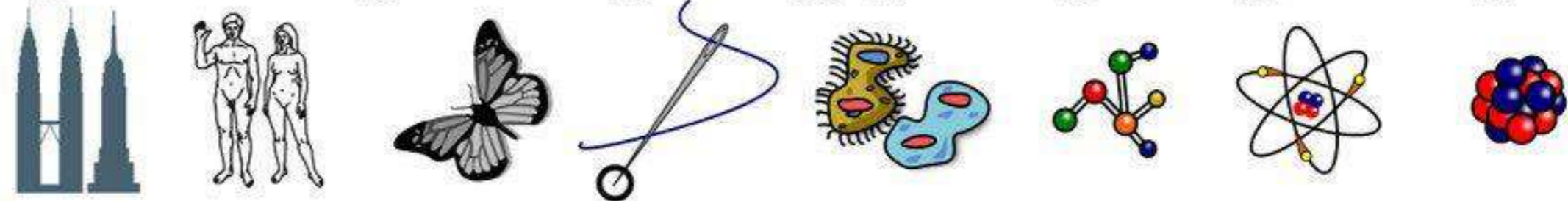
Penetrates Earth's Atmosphere?



Radiation Type
Wavelength (m)

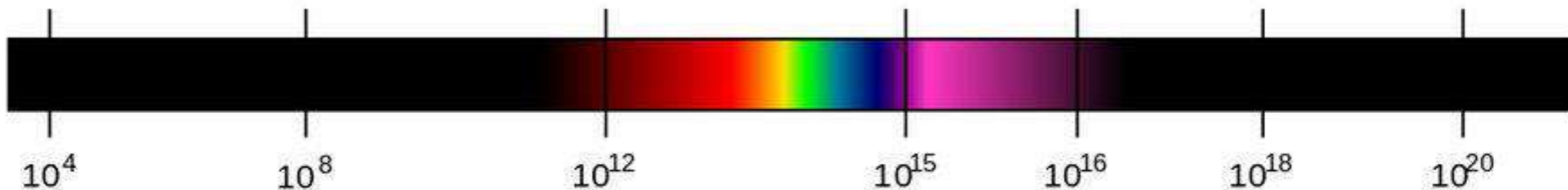


Approximate Scale
of Wavelength



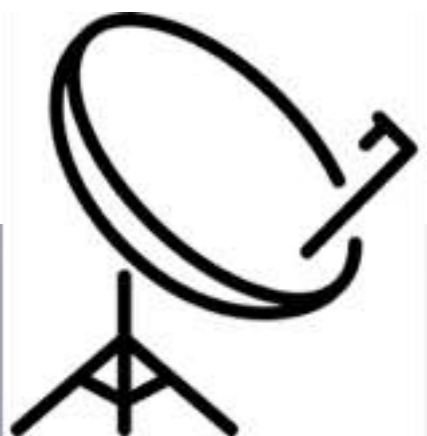
Buildings Humans Butterflies Needle Point Protozoans Molecules Atoms Atomic Nuclei

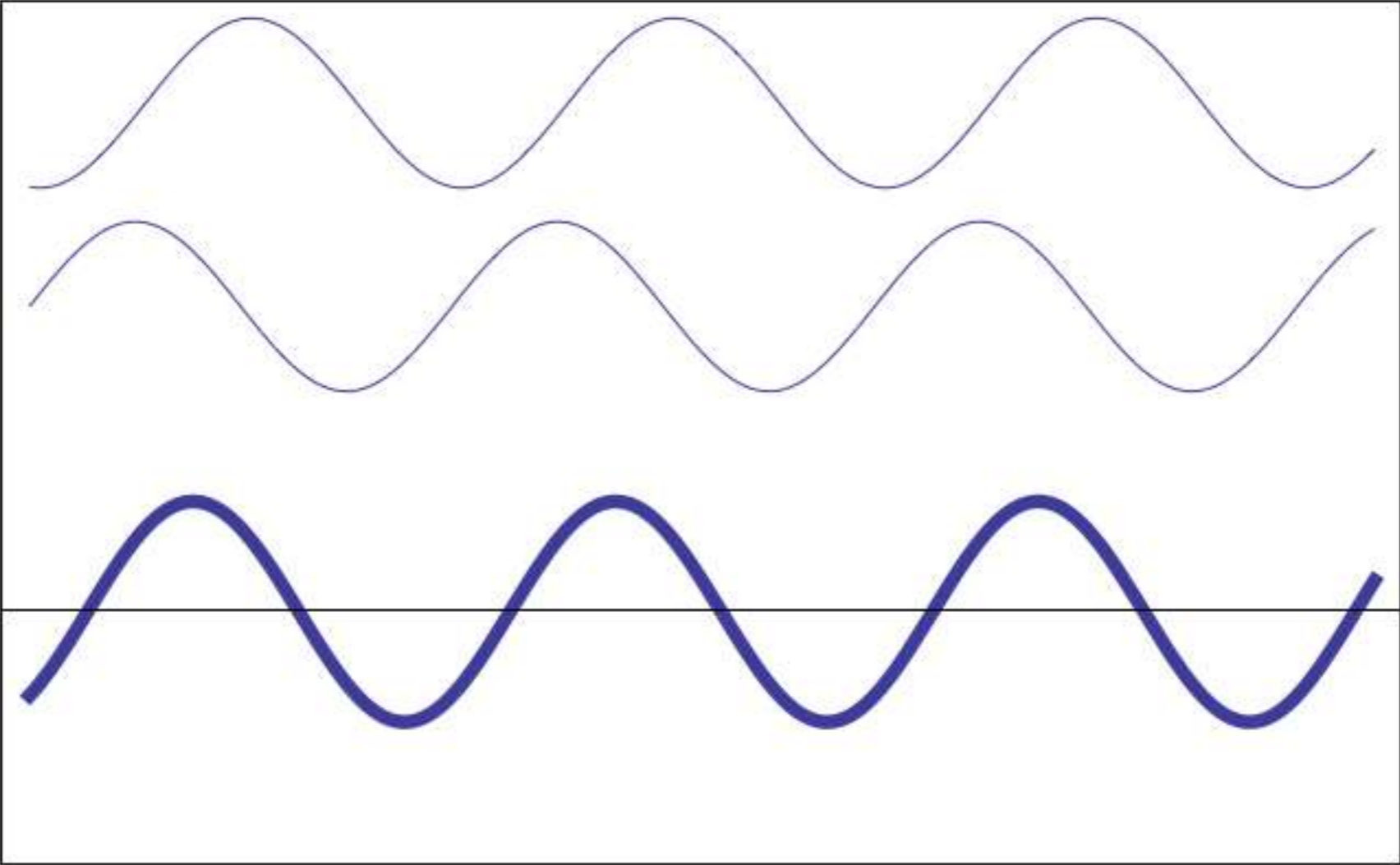
Frequency (Hz)



Interferometria

$$\text{Szögfelbontás} \sim \frac{\text{hullámhossz}}{\text{bázisvonalhossz}}$$





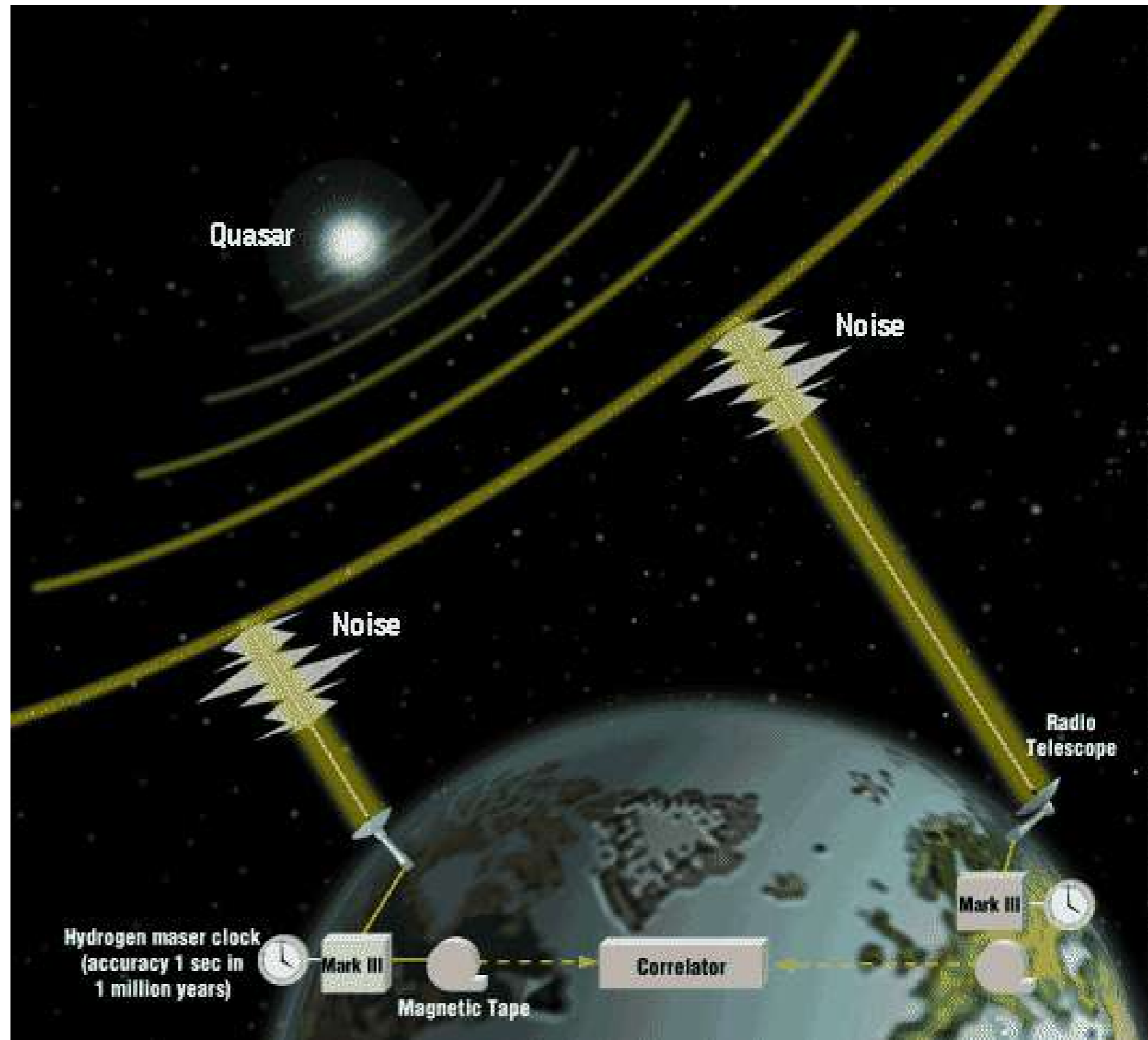
Nagyon hosszú bázisvonalú interferometria

Very Long Baseline Interferometry

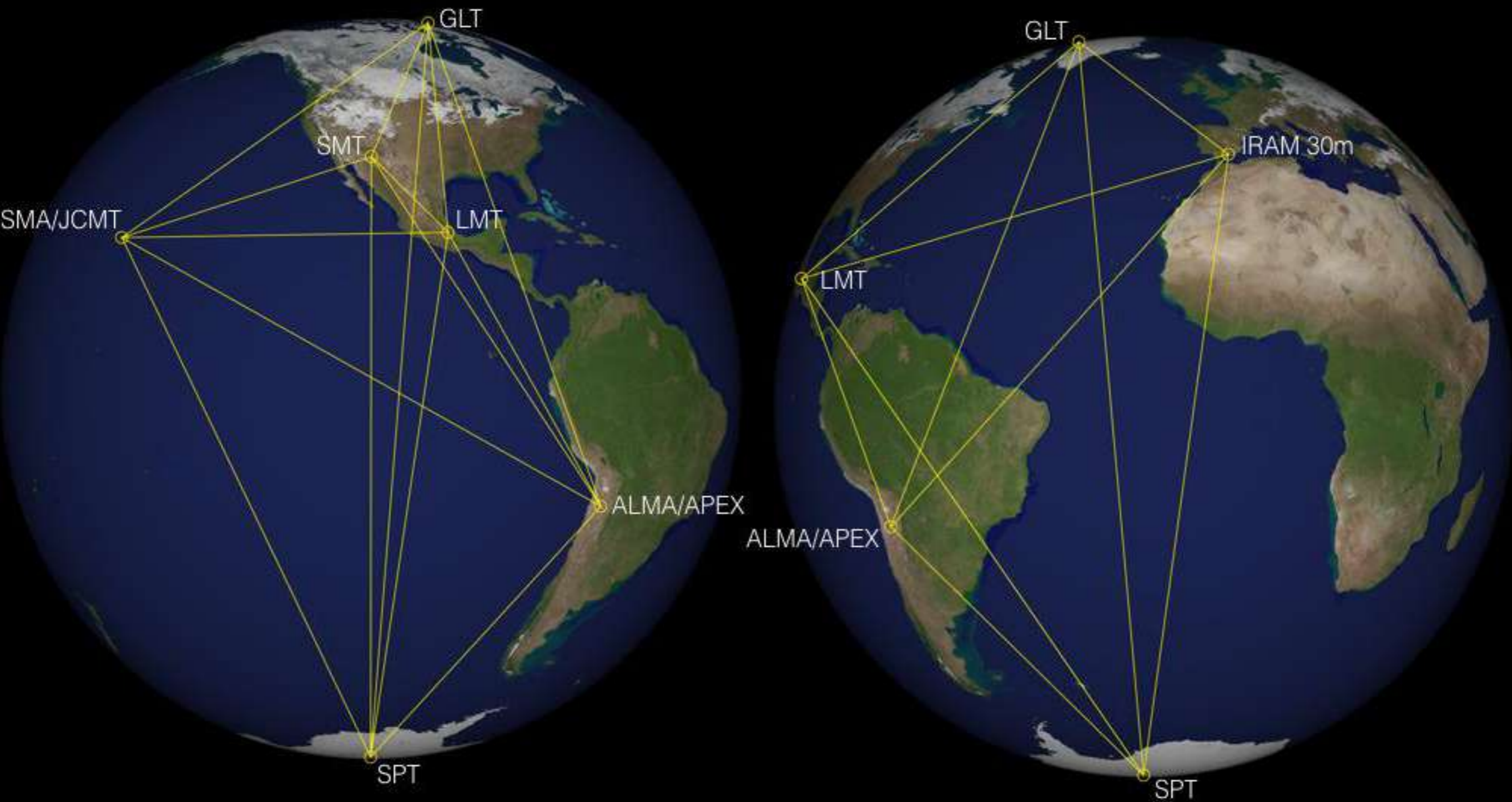
$$\theta \sim \frac{\lambda}{b}$$

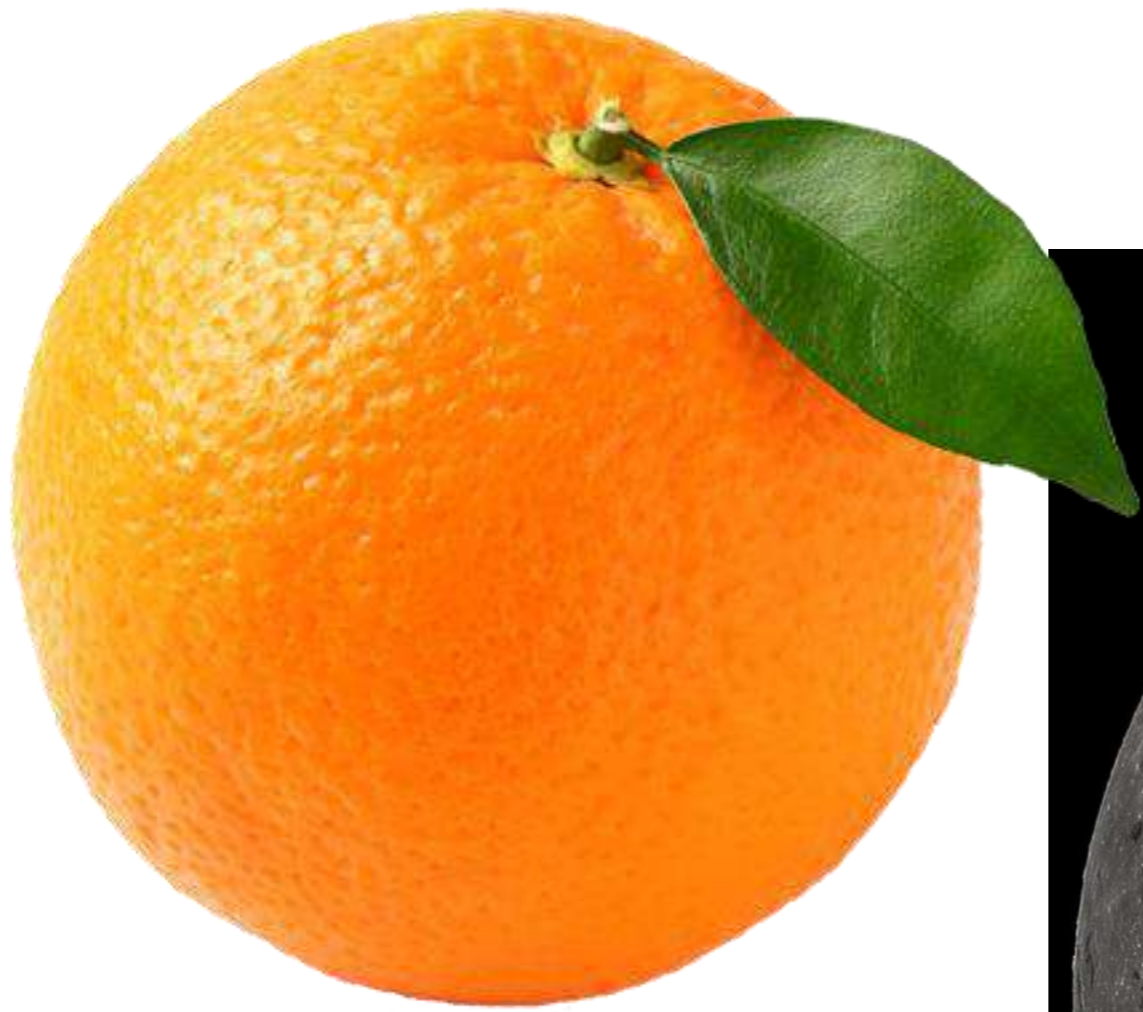
Hosszabb bázisvonal (b) \Rightarrow
finomabb szögfelbontás

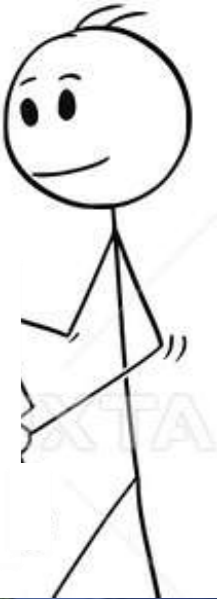
Rövidebb hullámhossz (λ)
 \Rightarrow finomabb
szögfelbontás



$$\lambda = 1.3 \text{ mm}, b = 10700 \text{ km} \Rightarrow \vartheta = 25 \mu\text{as}$$



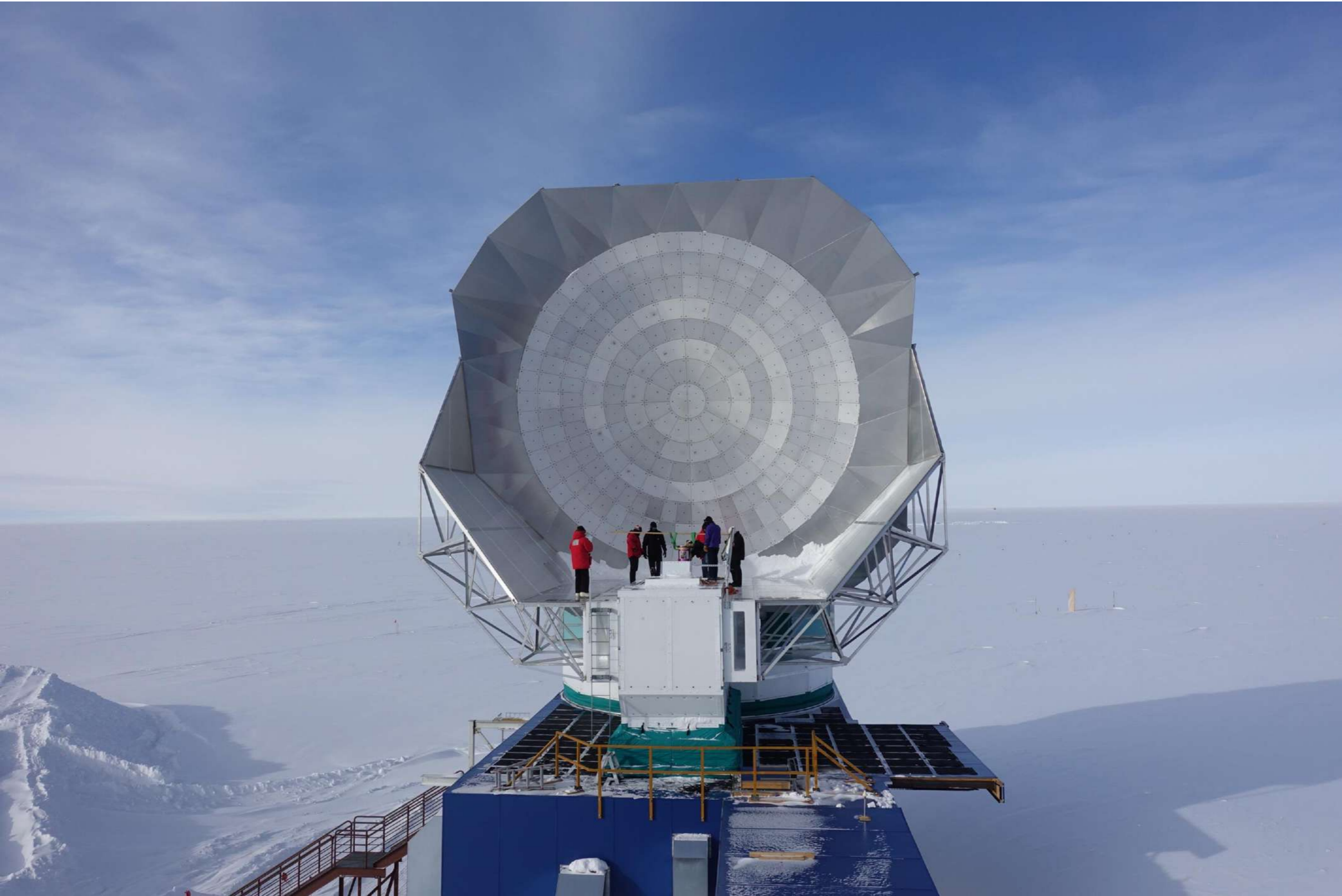




ALMA – Atacama Large Millimeter/Submillimeter Array



SPT – South Pole Telescope



LMT – Large Millimeter Telescope



Large Millimeter Telescope (Mexico)
által mért adatok fele, $\frac{1}{2}$ petabyte

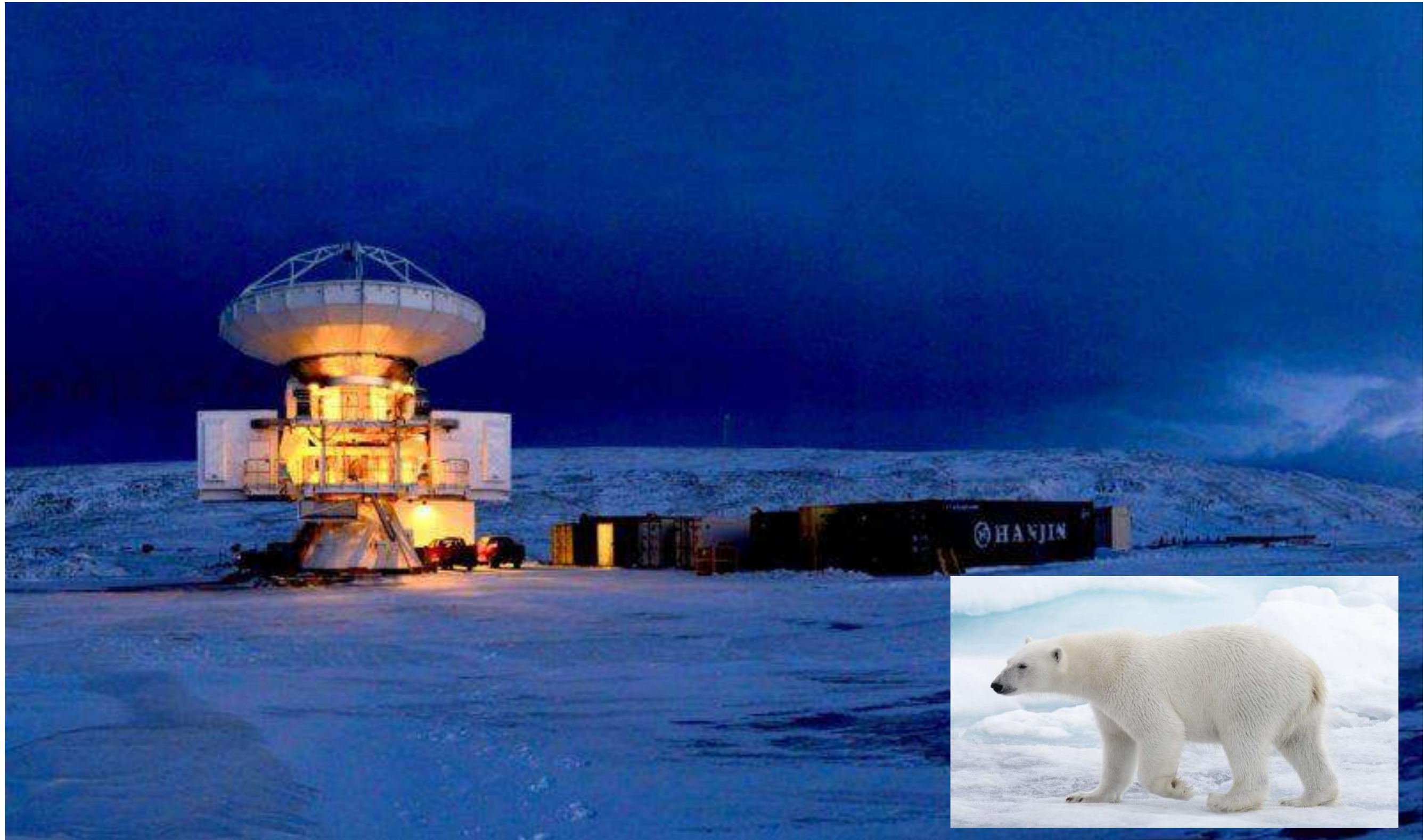
= Egy emberöltő alatt
készített szelfik mennyisége
5000 ember esetén



IRAM – Pico Veleta (Spanyolország)



GLT – Greenland Telescope (várhatóan 2021-től kapcsolódik be a mérésekbe)

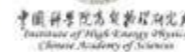


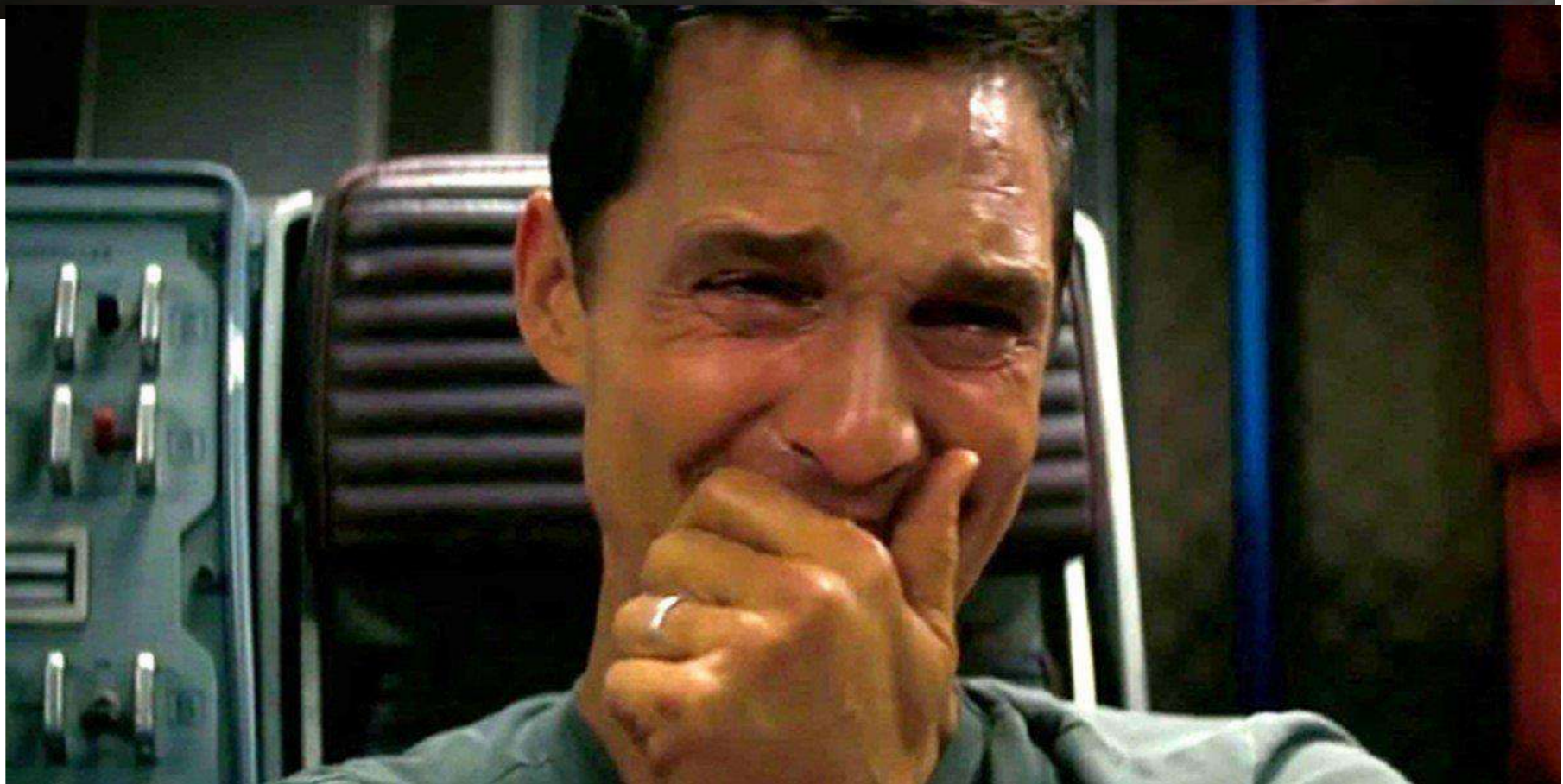


Large Millimeter Telescope *Alfonso Serrano*



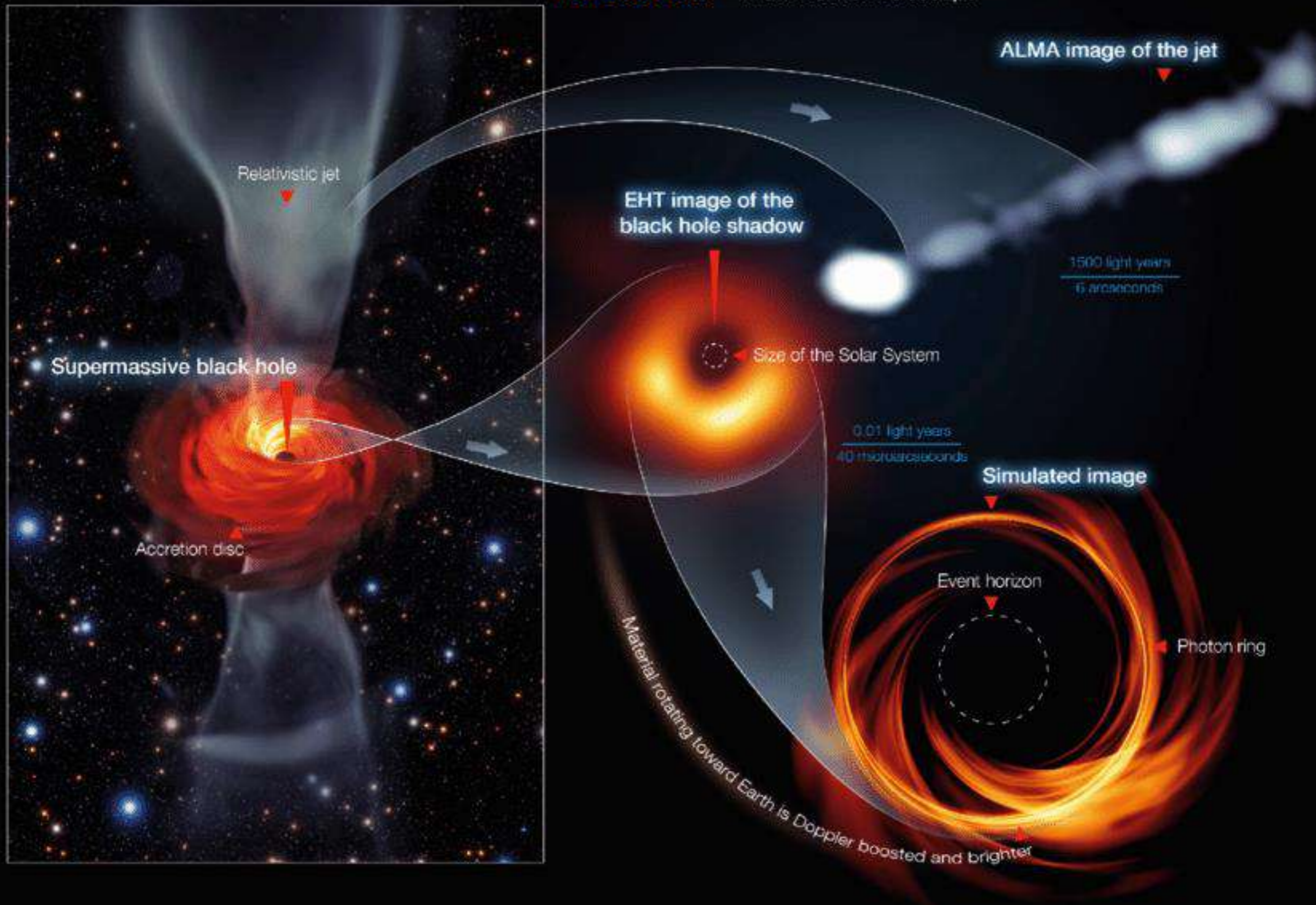
広島大学







M87 Black Hole – Event Horizon Telescope



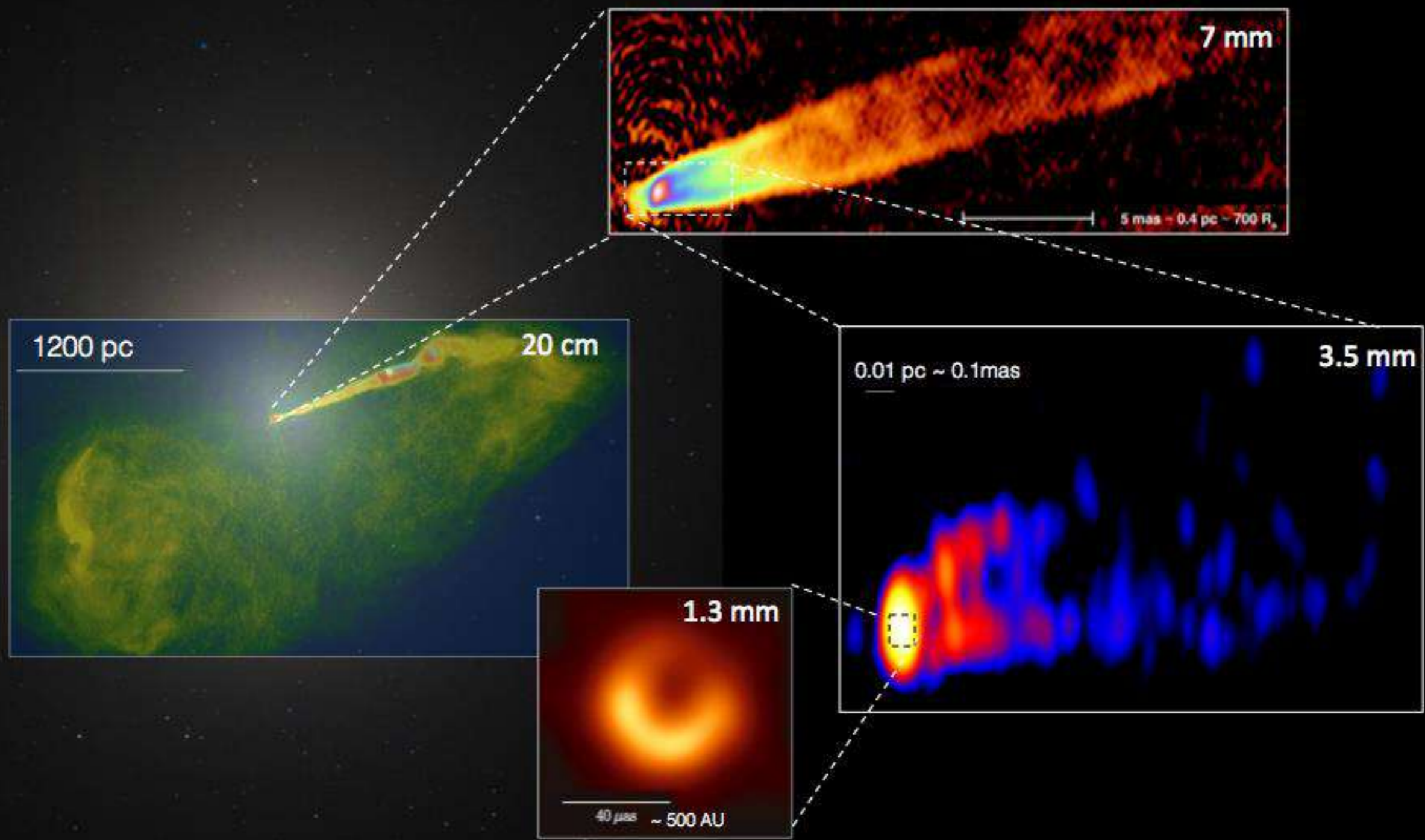
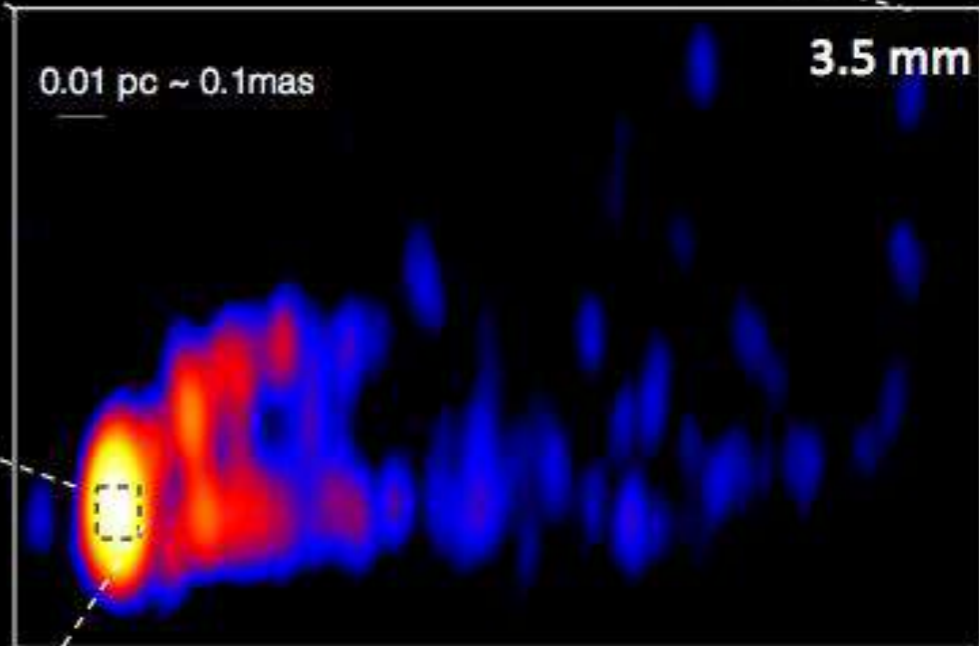
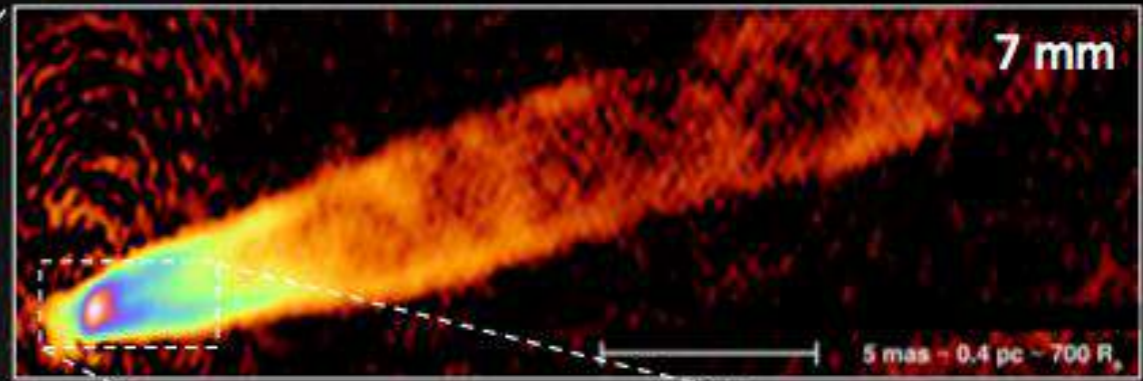
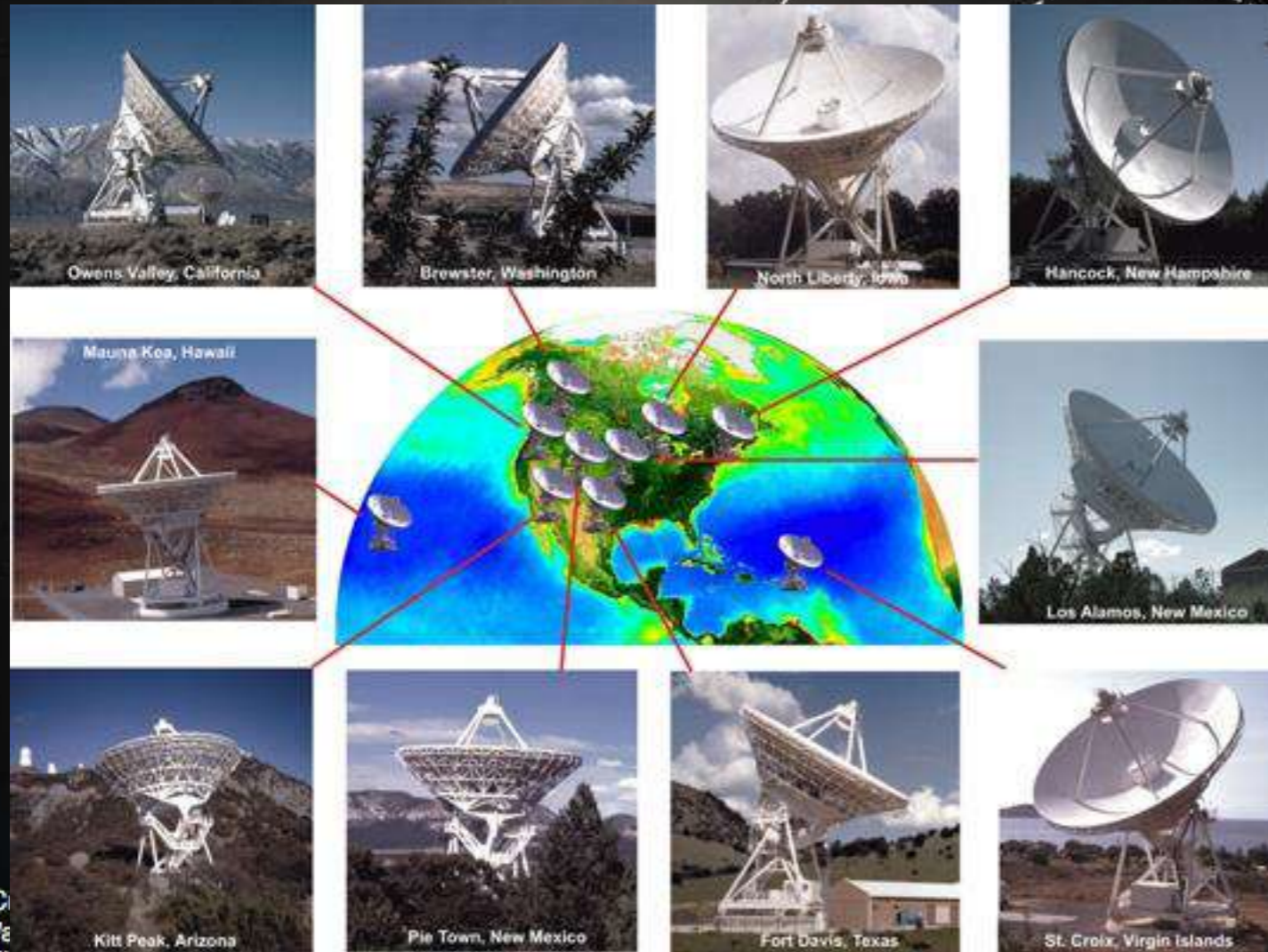


Image Credits: NASA/HST (optical), NRAO (VLA),
Craig Walker (7mm VLBA), Kazuhiro Hada (VLBA+GBT 3mm),
EHT Collaboration (1.3 mm)

Karl G. Jansky Very Large Array



Very Long Baseline Array



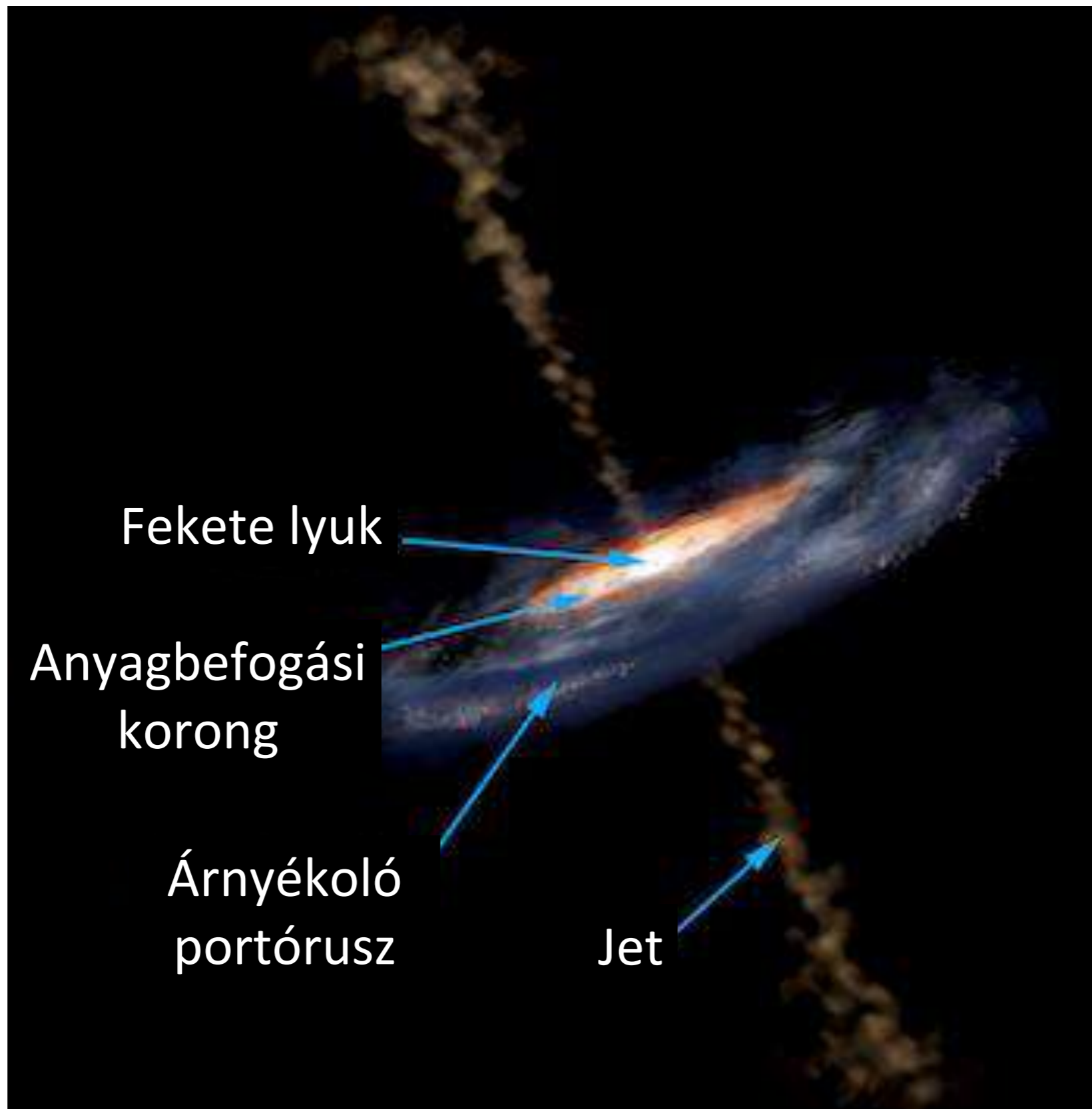


JIVE

Joint Institute for VLBI
ERIC



Image by Paul Boven (boven@jive.eu). Satellite image: Blue Marble Next Generation, courtesy of Nasa Visible Earth (visibleearth.nasa.gov).



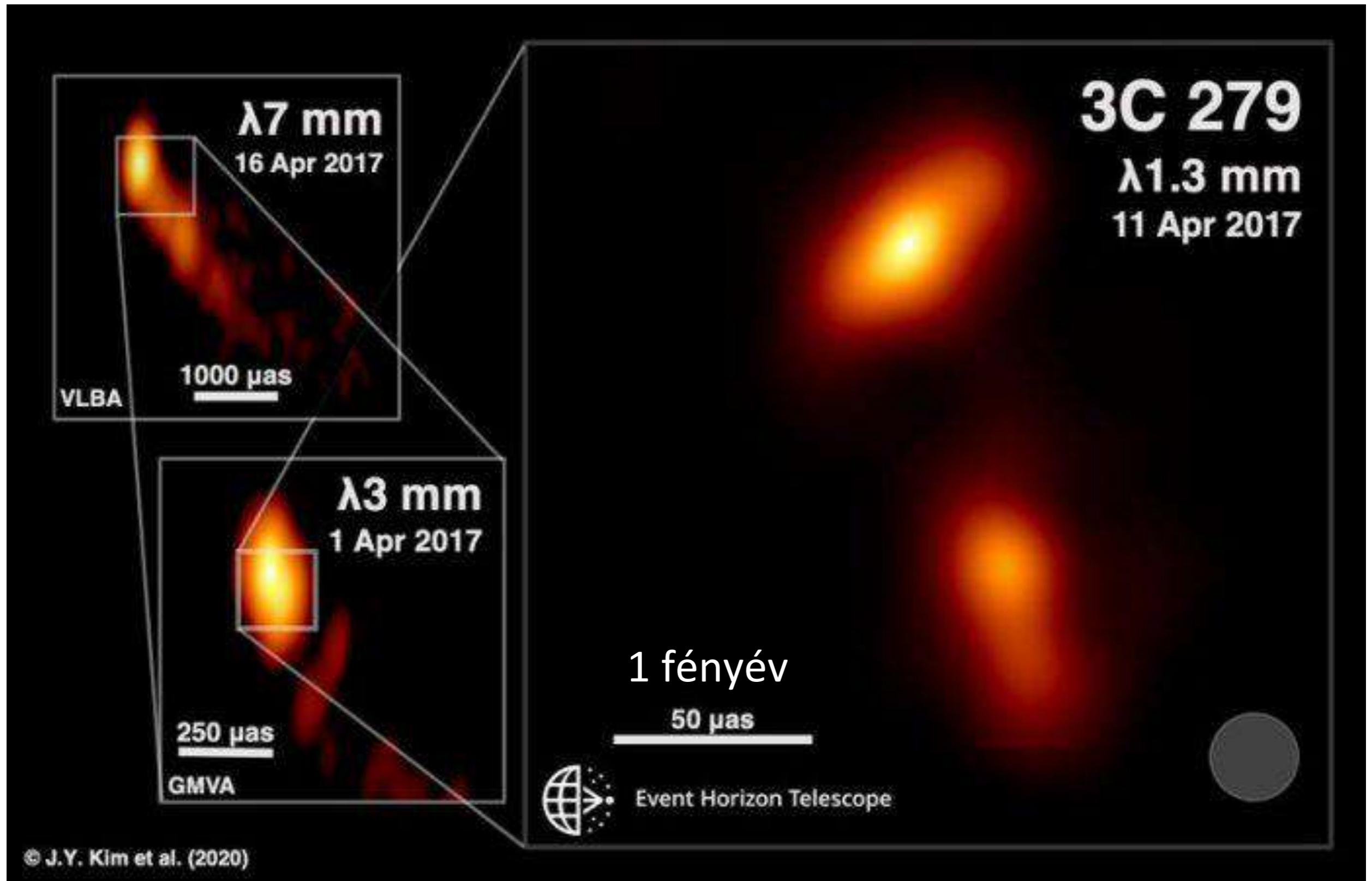
Fekete lyuk

Anyagbefogási
korong

Árnyékoló
portórusz

Jet

3C 279 – 5 milliárd fényévre



EHT-val érhető el a legfinomabb szögfelbontás?

Nem!

EHT-val érhető el a legfinomabb szögfelbontás?

Nem!

Űr-VLBI 2011-2019

Russia's RadioAstron space observatory

The RadioAstron observatory with an unprecedented high resolution capability will make it possible to observe remote objects in space

Parabolic antenna
• Diameter: 10 meters
• Comprises 27 carbon-plastic "petals"

Broad-beam antennas
Focal module

This is the first Russian orbital radio telescope

- It will study:
- Galaxy nuclei
 - Black holes
 - Neutron stars
 - Interstellar plasma clouds
 - The Earth's gravitational field
 - And many other objects and phenomena in the Universe

Ordered by: Federal Space Agency
Chief contractor: Lavochkin Research and Production Association

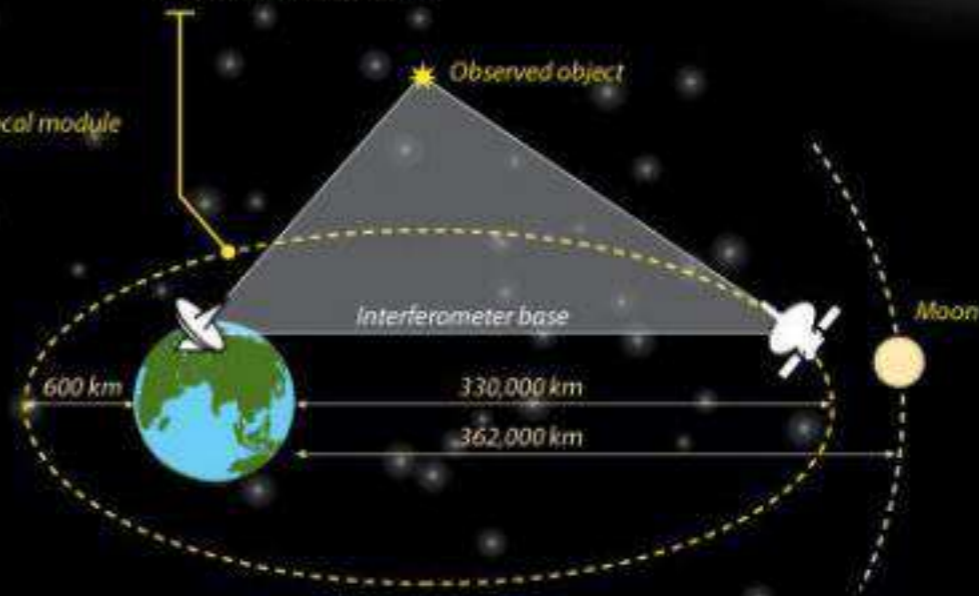
Scientific equipment developed by: Astro Space Center of the Russian Academy of Sciences' Lebedev Physics Institute

The RadioAstron observatory was launched on July 18, 2011.

Active service life: At least five years

Highly elliptical orbit

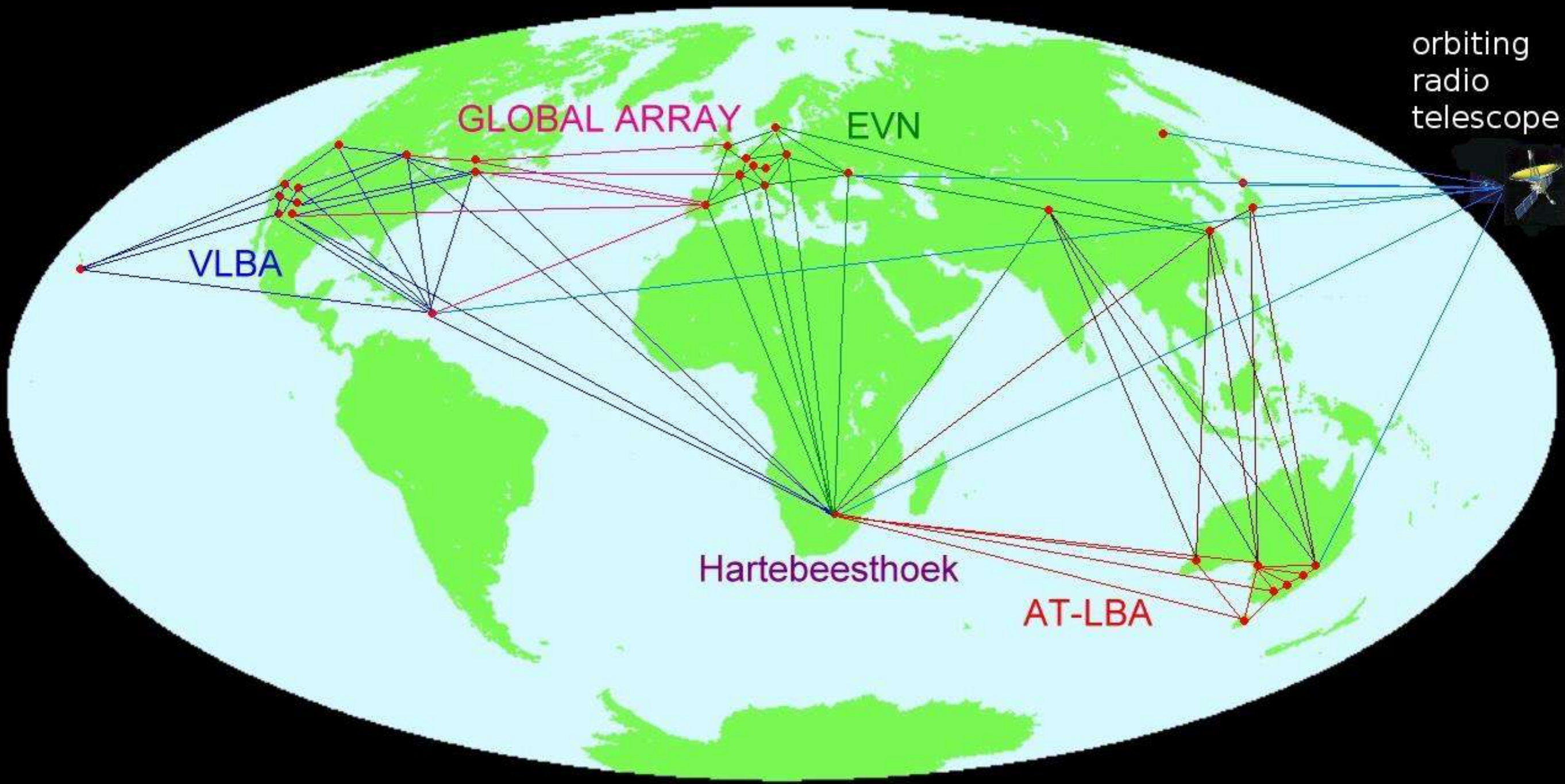
- Apogee: 330,000 kilometers
- Perigee: 600 km
- Orbital period: 8.2 days

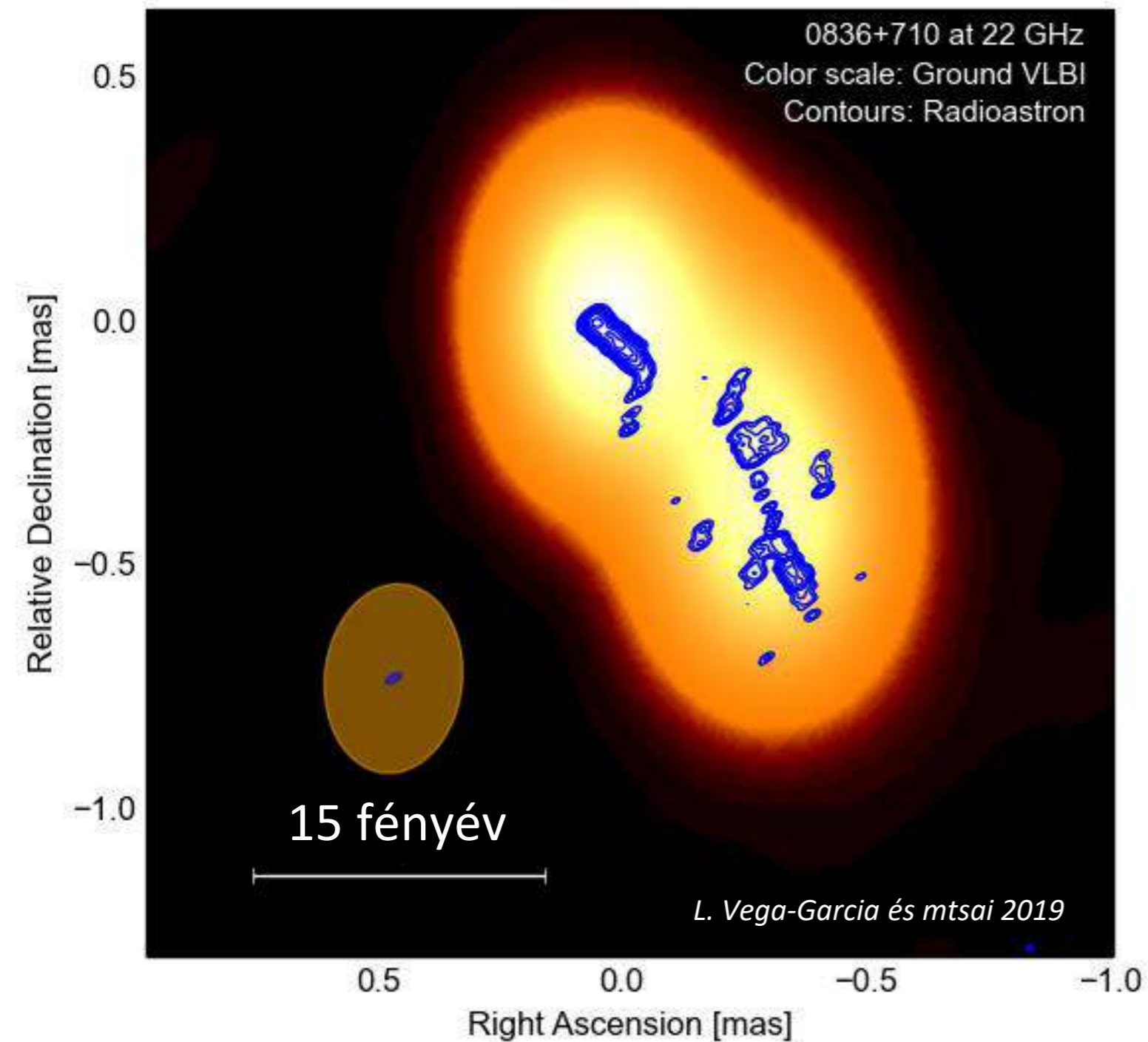


The RadioAstron observatory will operate with an international network of ground-based radio telescopes. This huge ground- and space-based telescope system, also called an interferometer, will provide the finest angular resolution.

This will make it possible to obtain images of remote objects with a resolution exceeding that of NASA's Hubble orbital telescope a thousand times over

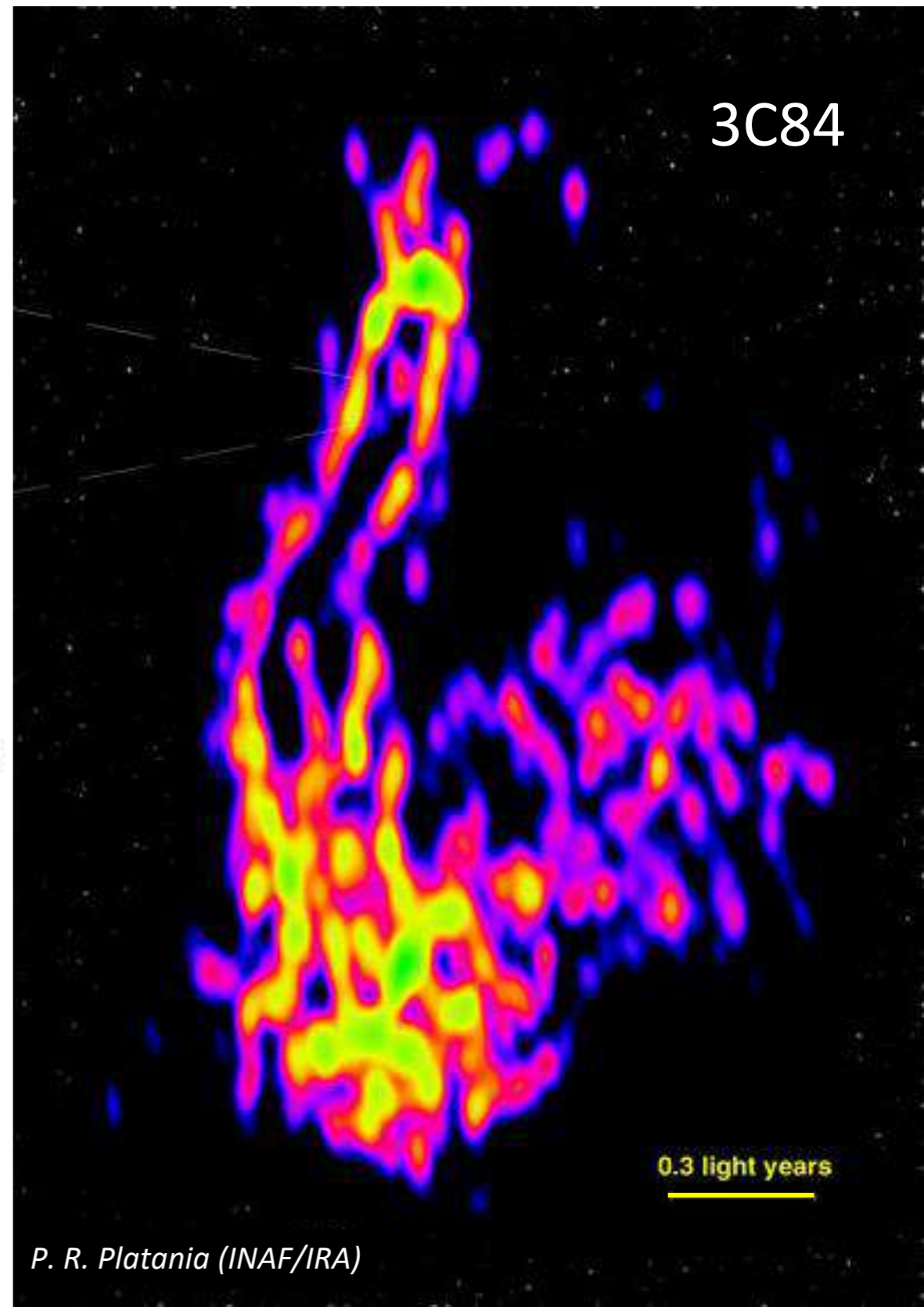
Radio Astronomy VLBI Arrays

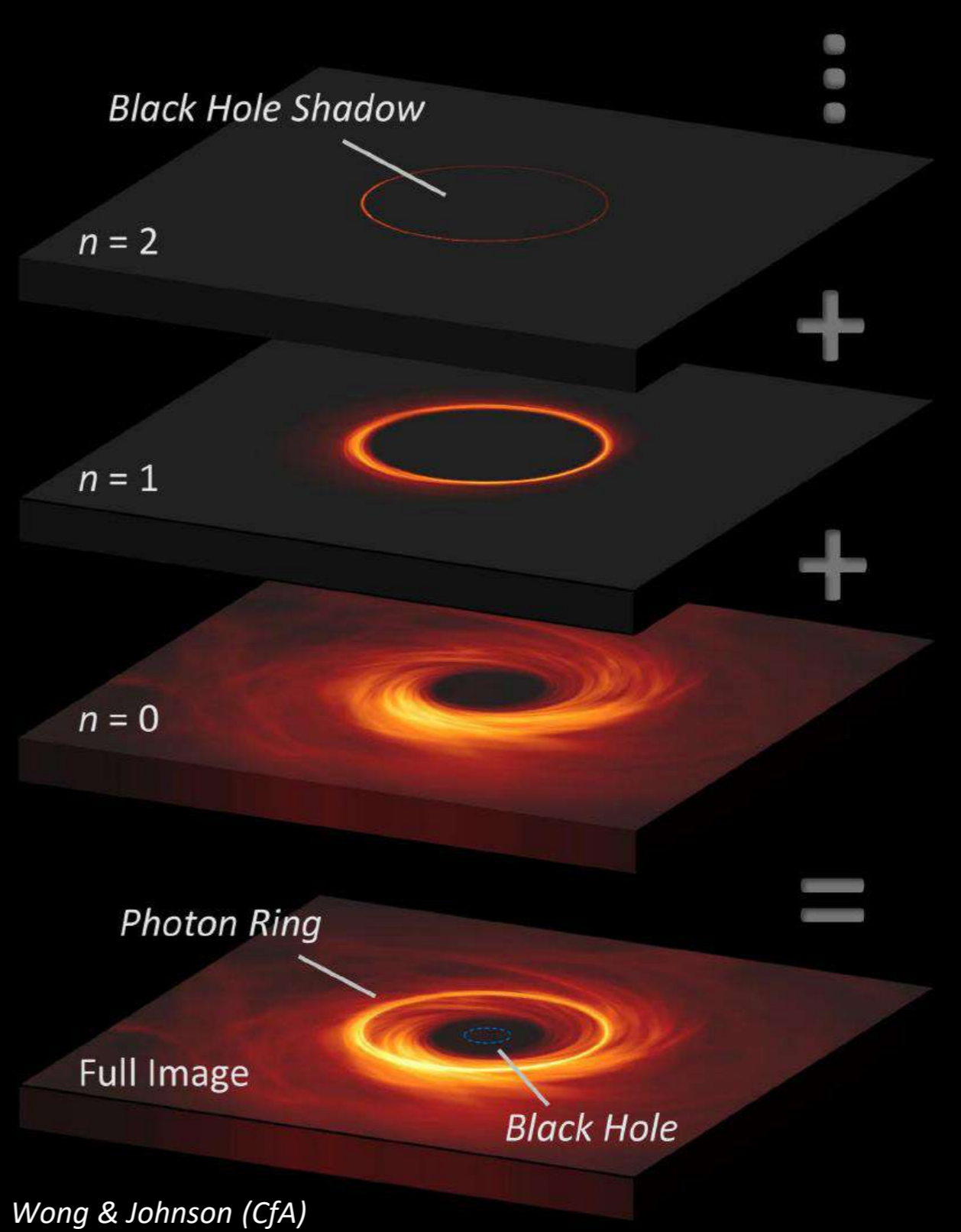
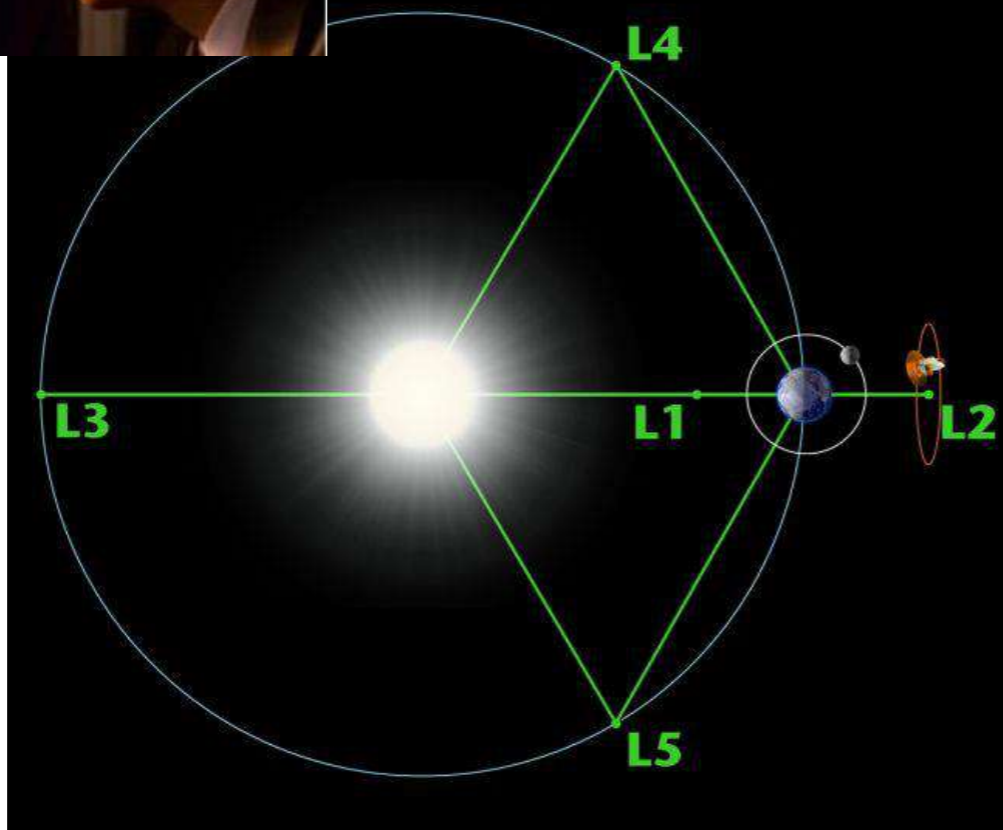
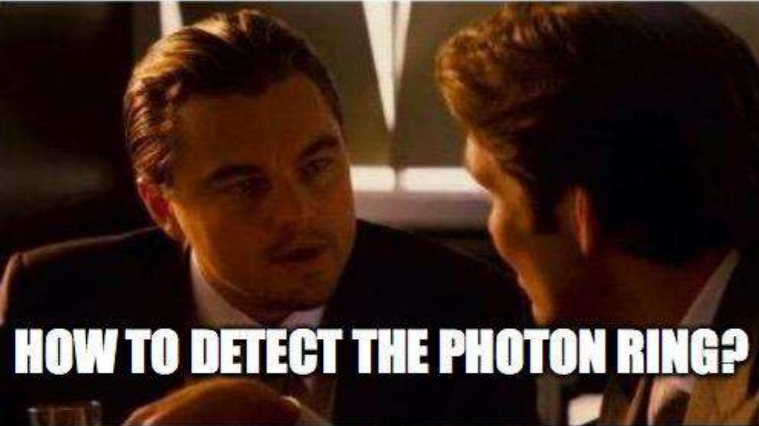




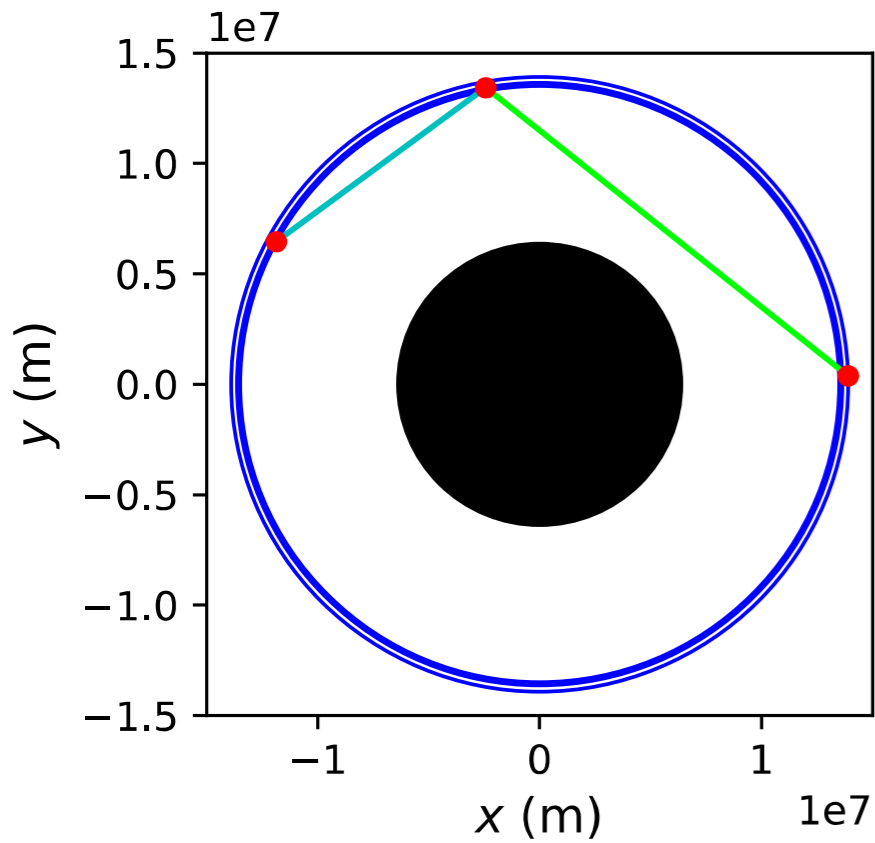
Távolság: 10 milliárd fényév

Távolság: 230 millió fényév



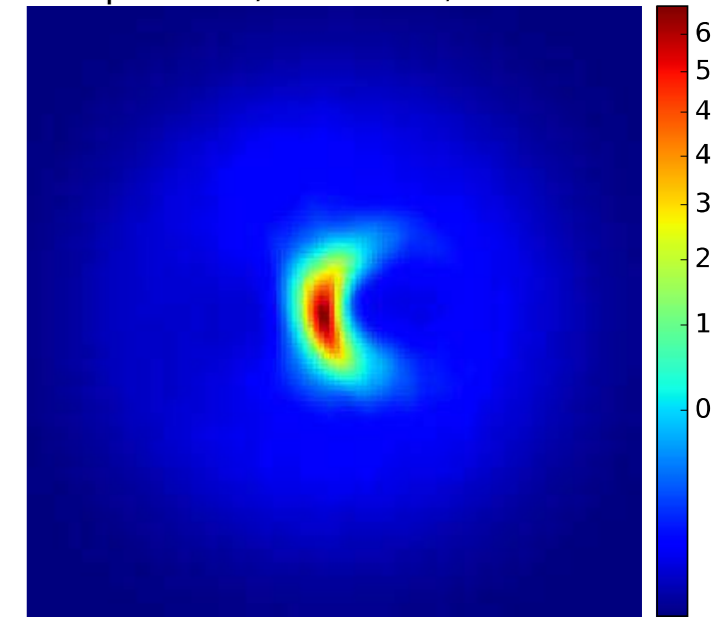


Javasolt űr-VLBI projektek mm-es hullámhosszon

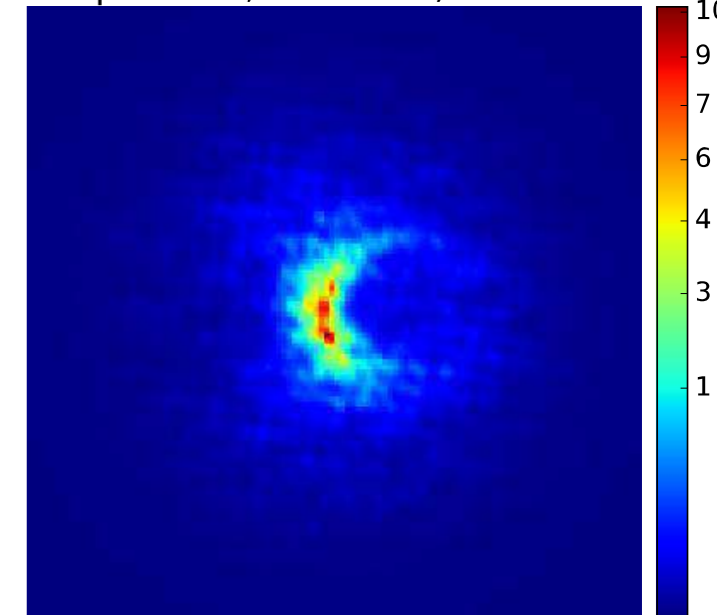


*Event Horizon
Imager: 3 keringő
antenna
Mérés 690 GHz-en*

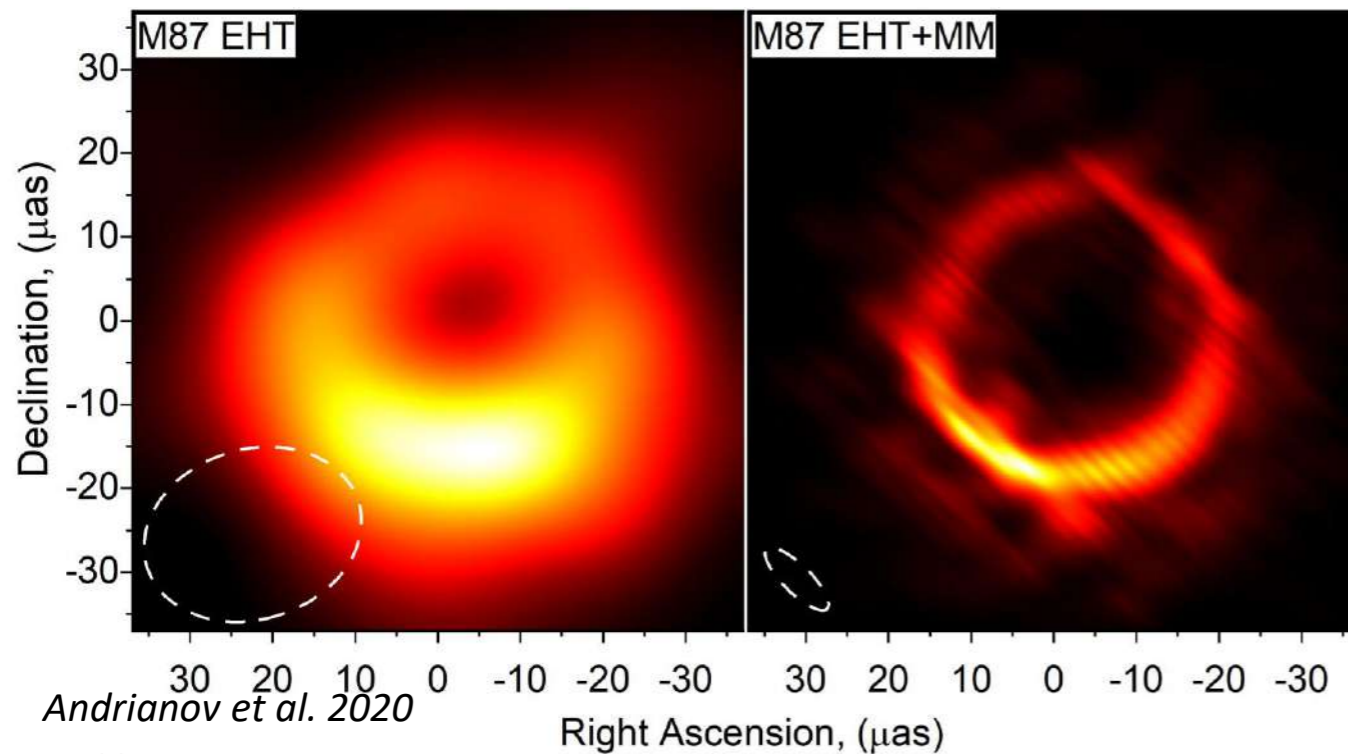
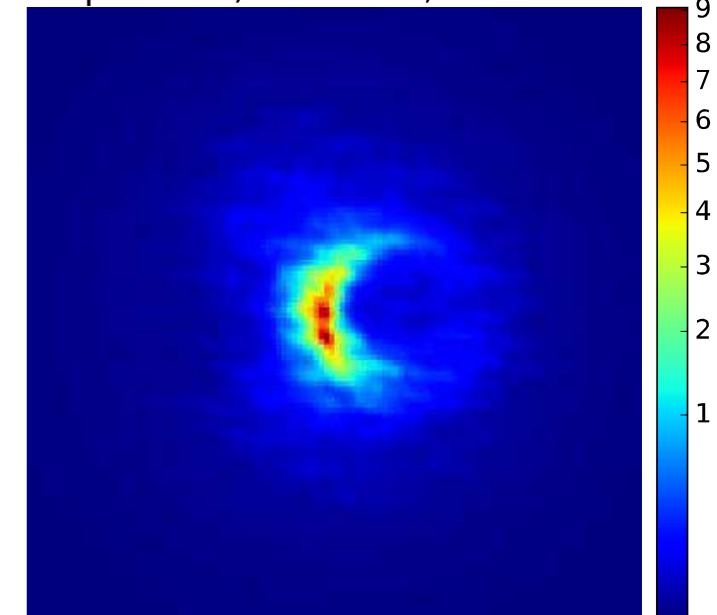
bispectrum, model 39, 1 month



bispectrum, model 39, 6 months

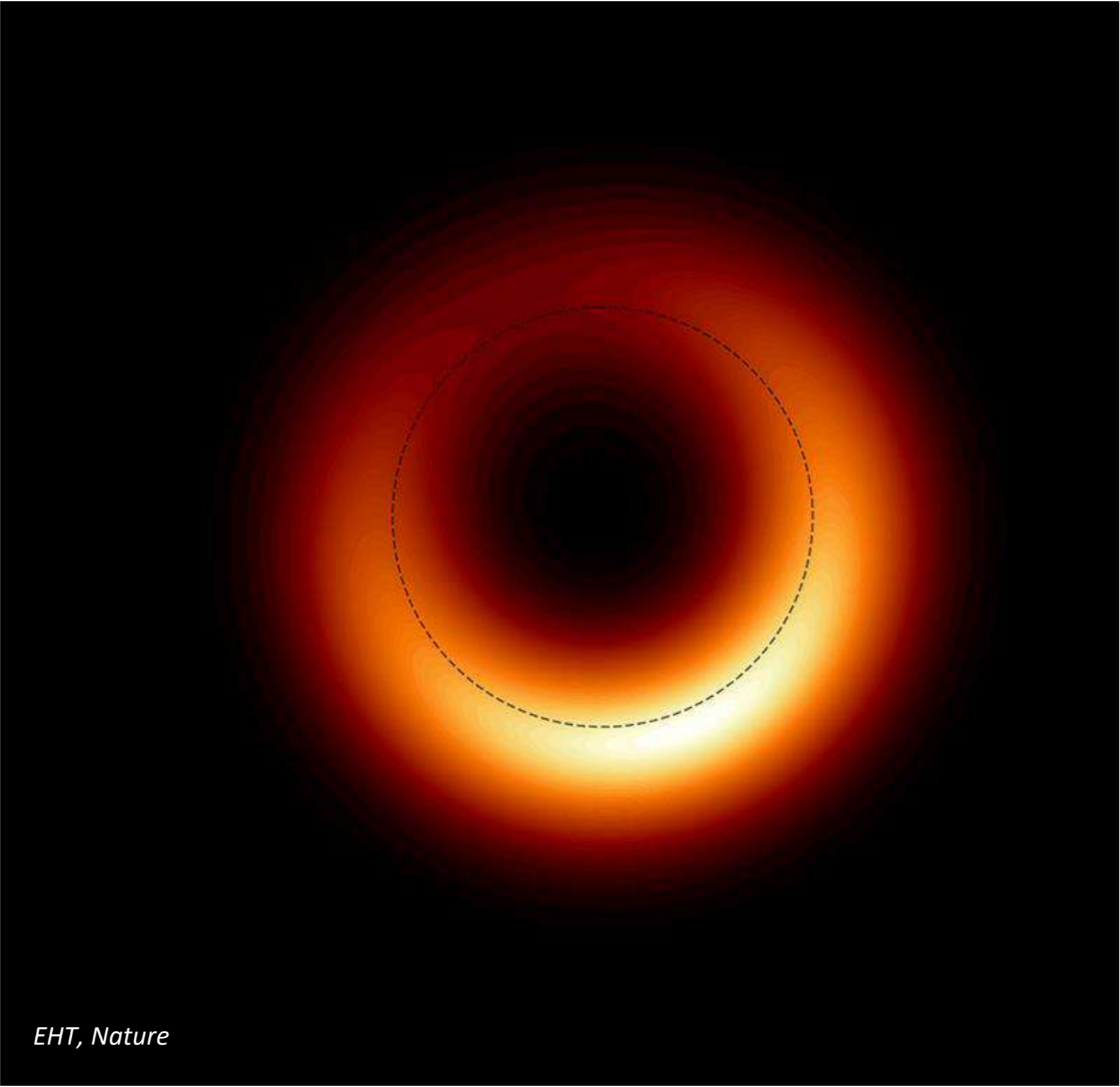


bispectrum, model 39, 24 months



Andrianov et al. 2020

Millimetron Space Observatory



EHT, Nature



Event Horizon Telescope

M 87*

2017

MODELS

2009

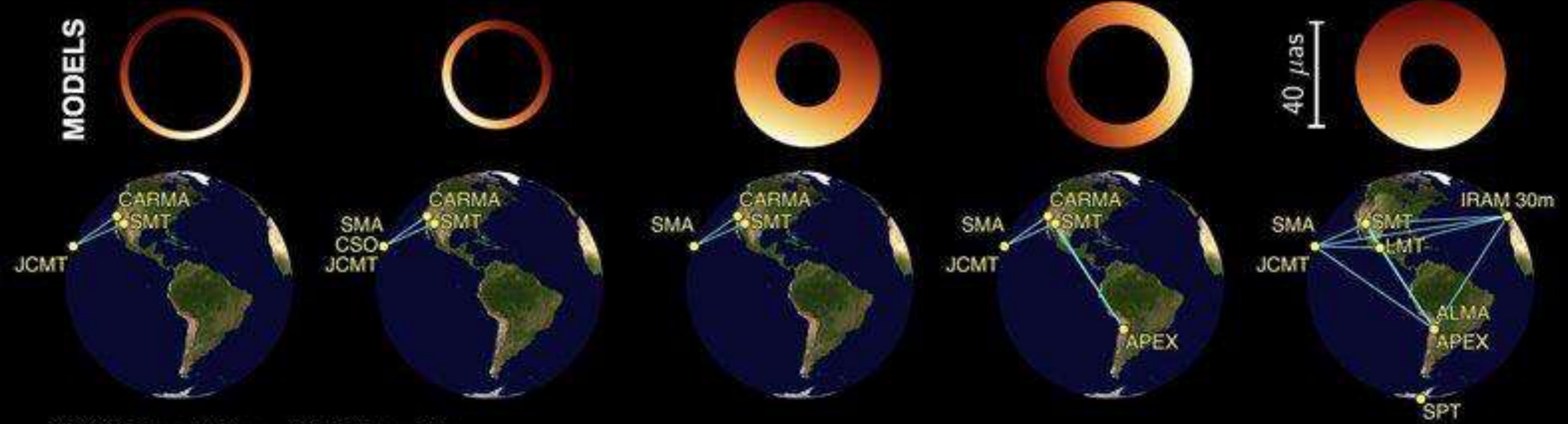
2011

2012

2013

IMAGE

40 μ as



© M. Wielgus, D. Pesce, EHT Collaboration